# SEPARATION OF VARIABLES AND LIE-ALGEBRA CONTRACTIONS. APPLICATIONS TO SPECIAL FUNCTIONS <br> G. Pogosyan 

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A review is given of some recently obtained results on analytic contractions of Lie algebras and Lie groups and their application to special function theory. The contractions considered are from $O(3)$ to $E(2)$ and from $O(2,1)$ to $E(2)$, or $E(1,1)$. The analytic contractions provide relations between separable coordinate systems on various homogeneous manifolds. They lead to asymptotic relations between basis functions and overlap functions for the representations of different groups.

## INTRODUCTION

Lie-algebra contractions were introduced into physics by Inönü and Wigner [1] in 1953 as a mathematical expression of a philosophical idea, namely the correspondence principle. This principle tells us that whenever a new physical theory surplants an old one, there should exist a well defined limit in which the results of the old theory are recovered. More specifically Inönü and Wigner established a relation between the Lorentz group and the Galilei one in which the former goes over into the latter as the speed of light satisfies $c \rightarrow \infty$.

The theory of Lie-algebra contractions (and deformations) has acquired a life of its own. It provides a framework in which large sets of Lie algebras can be embedded into families depending on parameters. All algebras in such a family have the same dimension, but they are not mutually isomorphic [2].

Two types of Lie-algebra contractions exist in the literature. The first are standard Inönü-Wigner contractions [1,3,4]. They can be interpreted as singular
changes of basis in a given Lie algebra $L$. Indeed, consider a basis $\left\{e_{1}, \ldots, e_{n}\right\}$ of $L$ and a transformation $f_{i}=U_{i k}(\varepsilon) e_{k}$, where the matrix $U$ realizing the transformation depends on some parameters $\varepsilon$. For $\varepsilon \rightarrow 0$ (i.e., some, or all of the components of $\varepsilon$ vanishing) the matrix $U(\varepsilon)$ is singular. In this limit the commutation relations of $L$ change (continuously) into those of a different, nonisomorphic, Lie algebra $L^{\prime}$.

More recently, «graded contractions» have been introduced [5-7]. They are more general than the Inönü-Wigner ones and can be obtained by introducing parameters modifying the structure constants of a Lie algebra $L_{1}$ in a manner respecting a certain grading and then taking limits when these parameters go to zero.

It is well known that there exists an intimate relationship between the theory of special functions and Lie group theory, well presented in the books of Vilenkin [8], Talman [9], and Miller [10]. In fact all properties of large classes of special functions can be obtained from the representation theory of Lie groups, making use of the fact that the special functions occur as basis functions of irreducible representations, as matrix elements of transformation matrices, as Clebsch-Gordon coefficients, or in some other guise. Recently, the class of functions treatable by group theoretical and algebraic methods has been extended to the so-called $q$-special functions that have been related to quantum groups [11-13].

One very fruitful application of Lie theory in this context is the algebraic approach to the separation of variables in partial differential equations [14-19]. In this approach separable coordinate systems (for Laplace-Beltrami, HamiltonJacobi and other invariant partial differential equations) are characterized by complete sets of commuting second order operators. These lie in the enveloping algebra of the Lie algebra of the isometry group, or in some cases of the conformal group, of the corresponding homogeneous space.

A question, that up to the last few years has received little attention in the literature, is that of connections between the separation of variables in different spaces, e.g., in homogeneous spaces of different Lie groups. In particular, it is of interest to study the behavior of separable coordinates, sets of commuting operators and the corresponding separated eigenfunctions under deformations and contractions of the underlying Lie algebras.

A recent series of papers [20-29] has been devoted to this new aspect of the theory of Lie-algebra and Lie-group contractions: the relation between the separation of variables in spaces of nonzero constant curvature and in flat spaces. The curved spaces were realized as spheres $S_{n} \sim O(n+1) / O(n)$, Lorentzian hyperboloids $H_{n} \sim O(n, 1) / O(n)$, or $O(n, 1) / O(n-1,1)$. The flat spaces where either Euclidean $E_{n}$, or pseudo-Euclidean $E(n-1,1)$ ones. The curved and flat spaces were related by a contraction of their isometry groups and the corresponding isotropy groups of the origin.

The essential point of these articles was the introduction of «analytic contractions». The contraction parameter is $R$ which is either the radius of the sphere
$S_{n}$, or the corresponding quantity $x_{0}^{2}-x_{1}^{2}-\ldots-x_{n}^{2}=R^{2}$ for the hyperboloid. The contractions are «analytic» because the parameter $R$ figures not only in the structure constants of the original Lie algebra, but also in the coordinate systems, in the operators of the Lie algebra, in the invariant operators characterizing the coordinate systems, in the separated eigenfunctions and the eigenvalues.

Once the parametrization displaying the parameter $R$ is established, it is possible to follow the contraction procedure $R \rightarrow \infty$ explicitly for all quantities: for the Lie algebra realized by vector fields, the Laplace-Beltrami operators, the second-order operators in the enveloping algebras, characterizing separable systems of coordinates, the separated ordinary differential equations, the eigenfunctions and the coefficients of the interbases expansions.

For two-dimensional spaces all types of coordinates were considered; for example, contractions of $O(3)$ to $E(2)$ relate elliptic coordinates on $S_{2}$ to elliptic and parabolic coordinates on $E_{2}$. They also relate spherical coordinates on $S_{2}$ to polar and Cartesian coordinates on $E_{2}$ [20,22,24]. Similarly, all 9 coordinate systems on the $H_{2}$ hyperboloid can be contracted to at least one of the four systems on $E_{2}$, or one of the 10 separable systems on $E_{1,1}$ [21,23].

Contractions from $S_{n}$ to $E_{n}$ were considered for subgroup type coordinates in Refs. 25, 26, 28, for subgroup type coordinates and certain types of elliptic and parabolic ones.

The main application of analytic contractions in this context is to derive special function identities, specially asymptotic formulas. Among other possible applications we mention the theory of finite dimensional integrable and superintegrable systems $[30,31]$.

In this paper we restrict ourselves to two-dimensional spaces of constant curvature.

## 1. SEPARATION OF VARIABLES IN TWO-DIMENSIONAL SPACES OF CONSTANT CURVATURE

1.1. Operator Approach to the Separation of Variables. Let us first consider a quite general two-dimensional Riemannian, pseudo-Riemannian or complex Riemannian spaces with a metric

$$
\begin{equation*}
d s^{2}=g_{i k} d u^{i} d u^{k}, \quad u=\left(u^{1}, u^{2}\right) \tag{1.1}
\end{equation*}
$$

In this space we introduce a classical free Hamiltonian

$$
\begin{equation*}
H=g_{i k}(u) p_{i} p_{k}, \tag{1.2}
\end{equation*}
$$

where $p_{i}=-\partial H / \partial u_{i}$ are the momenta classically conjugate to the coordinates $u^{i}$. We also introduce the corresponding Laplace-Beltrami operator

$$
\begin{equation*}
\Delta_{\mathrm{LB}}=\frac{1}{\sqrt{g}} \frac{\partial}{\partial u^{i}} \sqrt{g} g^{i k} \frac{\partial}{\partial u^{k}} \tag{1.3}
\end{equation*}
$$

We will be interested in two related question:

1. What are the quadratic polynomials on phase space

$$
\begin{equation*}
Q=a^{i k}(u) p_{i} p_{k} \tag{1.4}
\end{equation*}
$$

that Poisson commute with the Hamiltonian

$$
\begin{equation*}
\{Q, H\}=\sum_{i=1}^{2}\left(\frac{\partial Q}{\partial u_{i}} \frac{\partial H}{\partial p_{i}}-\frac{\partial Q}{\partial p_{i}} \frac{\partial H}{\partial u_{i}}\right)=0 ? \tag{1.5}
\end{equation*}
$$

In other words, when do quadratic (in the momenta) integrals of motion exist? Respectively what are the second-order Hermitian operators

$$
\begin{equation*}
Q=\left\{a^{i k}(u) \partial_{u_{i}} \partial_{u_{k}}\right\} \tag{1.6}
\end{equation*}
$$

(where the bracket denotes symmetrization) that Lie commute with the LaplaceBeltrami operator

$$
\begin{equation*}
[Q, H]=Q H-H Q=0 ? \tag{1.7}
\end{equation*}
$$

2. Do the Hamilton-Jacobi and Laplace-Beltrami equations in the considered space allow the separation of variables, and if so, how do we classify and construct separable coordinates? By separation of variables for the Hamilton-Jacobi equation we mean additive separation

$$
\begin{gather*}
g_{i k} \frac{\partial S}{\partial_{u^{i}}} \frac{\partial S}{\partial_{u^{k}}}=\lambda  \tag{1.8}\\
S=S_{1}\left(u^{1}, \lambda, \mu\right)+S_{2}\left(u^{2}, \lambda, \mu\right) \tag{1.9}
\end{gather*}
$$

For the Laplace-Beltrami operator we have in mind multiplicative separation

$$
\begin{equation*}
\Delta \Psi=\lambda \Psi, \quad \Psi=\Psi_{1}\left(u^{1}, \lambda, \mu\right) \Psi_{2}\left(u^{2}, \lambda, \mu\right) \tag{1.10}
\end{equation*}
$$

In both cases $\lambda$ and $\mu$ are the separation constants.
In this review article we shall mainly be interested in the Laplace-Beltrami operators in different spaces. However, some aspects of separation are simpler to discuss for the Hamilton-Jacobi equation. In two-dimensional Riemannian space the two equations separate in the same coordinate systems.

The existence of integrals of motion that are either linear or quadratic in the momenta was analyzed by Darboux [32] and Koenigs [33] in a note published in Volume 4 of Darboux's lectures. In particular it was shown that a metric (1.1) can allow $0,1,2,3$ or 5 linearly independent quadratic integrals. The case of 5 quadratic integrals occurs if and only if the metric corresponds to a space of constant curvature. In this case the second-order integrals are «reducible». That means that the metric allows precisely three first-order integrals

$$
\begin{equation*}
L_{i}=a_{i}(u) p_{1}+b_{i}(u) p_{2}, \quad u=\left(u_{1}, u_{2}\right), \quad i=1,2,3 \tag{1.11}
\end{equation*}
$$

and all second-order integrals are expressed as second-order polynomials (with constant coefficients) in terms of the first-order ones:

$$
\begin{equation*}
Q=\sum_{i, k=1}^{3} A_{i k} L_{i} L_{k}, \quad A_{i k}=\text { const. } \tag{1.12}
\end{equation*}
$$

If the polynomial $Q$ is the square of a first-order operator $L$, then it will provide a subgroup type coordinate. This is best seen by considering the corresponding first-order operator

$$
\begin{equation*}
X=\xi\left(u_{1}, u_{2}\right) \partial_{u_{1}}+\eta\left(u_{1}, u_{2}\right) \partial_{u_{1}} \tag{1.13}
\end{equation*}
$$

that generates a one-dimensional subgroup of the isometry group $G$. From $\left(u_{1}, u_{2}\right)$ we transform to the new coordinates $\left(v_{1}, v_{2}\right)$ «straightening out» the vector field (1.13) to the form

$$
\begin{equation*}
X=\partial_{v_{1}} \tag{1.14}
\end{equation*}
$$

Then $v_{1}$ will be an ignorable variable. The complementary variable $v_{2}=$ $\phi\left(u_{1}, u_{2}\right)$ can be replaced by an arbitrary function of $v_{2}$, the ignorable variable $v_{1}$ can be replaced by $f\left(v_{1}\right)+g\left(v_{2}\right)$, where both $f$ and $g$ are arbitrary. The separable coordinates are $v_{1}$ and $v_{2}$ (with the above-mentioned arbitrariness).

Now let us assume that an irreducible quadratic integral $Q$ as in (1.4) is known for a considered metric (1.1) (that is, $Q$ is not square of a linear integral). We can then impose that two equations be satisfied simultaneously. In the classical case they are

$$
\begin{equation*}
g_{i k} \frac{\partial S}{\partial_{u^{i}}} \frac{\partial S}{\partial_{u^{k}}}=\lambda, \quad a_{i k} \frac{\partial S}{\partial_{u^{i}}} \frac{\partial S}{\partial_{u^{k}}}=\mu \tag{1.15}
\end{equation*}
$$

Similarly, we can consider the quantum mechanics of a free particle in such a
space and write two simultaneous equations

$$
\begin{align*}
\hat{H} \Psi & =\left(\frac{1}{\sqrt{g}} \frac{\partial}{\partial u^{i}} \sqrt{g} g^{i k} \frac{\partial}{\partial u^{k}}\right) \Psi=\lambda \Psi,  \tag{1.16}\\
\hat{Q} \Psi & =\left(\frac{1}{\sqrt{g}} \frac{\partial}{\partial u^{i}} \sqrt{g} a^{i k} \frac{\partial}{\partial u^{k}}\right) \Psi=\mu \Psi .
\end{align*}
$$

Separable coordinates for the two systems (1.15) and (1.16) are obtained by simultaneously transforming $\hat{H}$ and $\hat{Q}$ to a standard form, in which the matrices $g_{i k}$ and $a_{i k}$ of (1.15) (and (1.16)) are diagonal. This can be done by solving the characteristic equation

$$
\begin{equation*}
\left|a_{i k}-\rho g_{i k}\right|=0 . \tag{1.17}
\end{equation*}
$$

If two distinct roots $\rho_{1}$ and $\rho_{2}$ exist, they will provide separable coordinates, at least over the field of complex number. If we are considering real spaces, then it may happen that $\rho_{1}$ and $\rho_{2}$ are real only in part of the space and do not parametrize the entire space. We will see below that this indeed happens for instance in the pseudo-Euclidean plane $E_{1,1}$.

The roots $\rho_{1}$ and $\rho_{2}$ can be replaced by any functions $u=u\left(\rho_{1}\right), v=v\left(\rho_{2}\right)$. This freedom can be used to transform $H$ and $Q$ simultaneously to the form

$$
\begin{gather*}
H=\frac{1}{\alpha(v)+\beta(u)}\left(p_{u}^{2}+p_{v}^{2}\right)=\lambda,  \tag{1.18}\\
Q=\frac{1}{\alpha(v)+\beta(u)}\left(\beta(u) p_{u}^{2}-\alpha(v) p_{v}^{2}\right)=\mu .
\end{gather*}
$$

The Hamiltonian $H$ in (1.18) is in its Liouville form [34]. The separated equations are

$$
\begin{equation*}
\alpha H+Q=\alpha \lambda+\mu, \quad \beta H-Q=\beta \lambda-\mu \tag{1.19}
\end{equation*}
$$

for the Hamilton-Jacobi equation and similarly

$$
\begin{equation*}
(\alpha \hat{H}+\hat{Q}) \Psi=(\alpha \lambda+\mu) \Psi, \quad(\beta \hat{H}-\hat{Q}) \Psi=(\beta \lambda-\mu) \Psi \tag{1.20}
\end{equation*}
$$

for the Laplace-Beltrami equation.
Let us now restrict ourselves to two-dimensional spaces $M$ of constant curvature, that is to the Euclidean plane $E_{2}$, pseudo-Euclidean plane $E_{1,1}$, sphere $S_{2}$ and two-, or one-sheeted hyperboloid $H_{2}$. Each of these has a three-dimensional isometry group $G$. The Lie algebra $L$ of $G$ has in each case a standard basis which we denote $\left\{X_{1}, X_{2}, X_{3}\right\}$.

The Laplace-Beltrami operator $\hat{H}=\Delta_{\mathrm{LB}}$ (1.3) is in each case proportional to the Casimir operator $\hat{C}$ of the Lie algebra $L$. The operator $Q$ commuting with $H$ will have the form

$$
\begin{equation*}
\hat{Q}=\sum_{i, k=1}^{3} A_{i k} X_{i} X_{k}, \quad A_{i k}=A_{k i} \tag{1.21}
\end{equation*}
$$

where $A$ is a constant matrix. Let $g \in G$ be an element of the Lie algebra of the isometry group of the considered space. Let us rewrite Eq. (1.21) in matrix form

$$
\begin{equation*}
\hat{Q}=u^{T} A u, \quad u^{T}=\left\{X_{1}, X_{2}, X_{3}\right\} \tag{1.22}
\end{equation*}
$$

The transformation $g$ acting on the space $M$ induces a transformation $u^{\prime}=g u$ on the Lie algebra $L$. The Casimir operator $\Delta_{\mathrm{LB}}$ stays invariant, but $\hat{Q}$ transforms to

$$
\begin{equation*}
\hat{Q}=u^{\prime T} g^{T} A g u^{\prime} \tag{1.23}
\end{equation*}
$$

Thus, for spaces of constant curvature a classification of operators $\hat{Q}$ commuting with $\hat{H}$ reduces to a classification of symmetric matrices $A=A^{T}$ into equivalence classes under the congruence transformation

$$
\begin{equation*}
A^{\prime}=g^{T} A g, \quad g \in G \tag{1.24}
\end{equation*}
$$

This problem, as we shall see below, can be reduced to that of classifying elements of Jordan algebras into equivalence classes.

Furthermore, the operator $\hat{C}$ of $L$ can also be written in the form

$$
\hat{C}=u^{T} C u, \quad C=\left(\begin{array}{ccc}
c_{1} & &  \tag{1.25}\\
& c_{2} & \\
& & c_{3}
\end{array}\right)
$$

Two matrices $A$ and $\tilde{A}$ will give equivalent coordinate systems if they satisfy

$$
\begin{equation*}
\tilde{A}=\lambda g^{T} A g+\mu C, \quad \lambda \neq 0 \tag{1.26}
\end{equation*}
$$

where $\lambda$ and $\mu$ are real constants.
1.2. Separable Coordinate Systems in the Euclidean Plane. The Lie algebra of the isometry group $E(2)$ is given by

$$
\begin{equation*}
L=x_{2} \partial_{x_{1}}-x_{1} \partial_{x_{2}}, \quad P_{1}=\partial_{x_{1}}, \quad P_{2}=\partial_{x_{2}} \tag{1.27}
\end{equation*}
$$

The operator $\hat{Q}$ of Eq. (1.21) will in this case be

$$
\begin{align*}
\hat{Q}=a L^{2}+b_{1}\left(L P_{1}+P_{1} L\right)+b_{2}\left(L P_{2}+\right. & \left.P_{2} L\right)+ \\
& +c_{1} P_{1}^{2}+c_{2} P_{2}^{2}+2 c_{3} P_{1} P_{2} \tag{1.28}
\end{align*}
$$

An $E(2)$ transformation matrix will be written as

$$
g=\left(\begin{array}{cc}
1 & \xi^{T}  \tag{1.29}\\
0 & R
\end{array}\right), \quad \xi^{T}=\left(\xi_{1}, \xi_{2}\right), \quad R \in \mathbb{R}^{2}, \quad R^{T} R=I
$$

The matrix $A$ of Eq. (1.22) is

$$
A=\left(\begin{array}{cc}
a & b^{T}  \tag{1.30}\\
b & S
\end{array}\right), \quad S=\left(\begin{array}{cc}
c_{1} & c_{3} \\
c_{3} & c_{2}
\end{array}\right), \quad b^{T}=\left(b_{1}, b_{2}\right)
$$

and $C$ of (1.25) is

$$
C=\left(\begin{array}{ccc}
0 & &  \tag{1.31}\\
& 1 & \\
& & 1
\end{array}\right)
$$

since the Casimir operator is

$$
\begin{equation*}
C=\Delta=P_{1}^{2}+P_{2}^{2} \tag{1.32}
\end{equation*}
$$

The transformation (1.26) with $\lambda=1$ has two invariants in the space of symmetric matrices $A$, namely

$$
\begin{equation*}
I_{1}=a, \quad I_{2}=\left\{\left[a\left(c_{1}-c_{2}\right)-\left(b_{1}^{2}-b_{2}^{2}\right)\right]^{2}+4\left(a c_{3}-b_{1} b_{2}\right)^{2}\right\}^{1 / 2} \tag{1.33}
\end{equation*}
$$

Correspondingly, the operator $\hat{Q}$ can be transformed into one of four canonical forms

1) $I_{1}=0, \quad I_{2}=0, \quad Q_{C}=P_{1}^{2}$,
2) $I_{1} \neq 0, \quad I_{2}=0, \quad Q_{R}=L^{2}$,
3) $I_{1}=0, \quad I_{2} \neq 0, \quad Q_{P}=L P_{2}+P_{2} L$,
4) $I_{1} \neq 0, \quad I_{2} \neq 0, \quad Q_{E}=L^{2}-D^{2} P_{2}^{2}, \quad D^{2}=\frac{I_{2}}{I_{1}^{2}}$.

The first two forms correspond to subgroup-type coordinates. Thus, $Q_{C}$ of (1.34) corresponds to Cartesian coordinates, in which $P_{1}=\partial_{x}$ (and also $P_{2}=\partial_{y}$ ) is already straightened out. Both $x$ and $y$ are ignorable. The second, $Q_{R}$, corresponds to polar coordinates

$$
\begin{equation*}
x=r \cos \phi, \quad y=r \sin \phi \tag{1.38}
\end{equation*}
$$

in which $L=\partial_{\phi}$ is straightened out so that $\phi$ is an ignorable variable.
The coordinates corresponding $Q_{P}$ of (1.36) are the parabolic coordinates

$$
\begin{equation*}
x=\frac{1}{2}\left(u^{2}-v^{2}\right), \quad y=u v \tag{1.39}
\end{equation*}
$$

Equivalently, if we take $\tilde{Q}_{P}=L P_{1}+P_{1} L$, the prescription (1.17) leads to

$$
\begin{equation*}
x=u v, \quad y=\frac{1}{2}\left(u^{2}-v^{2}\right) . \tag{1.40}
\end{equation*}
$$

Finally, $Q_{E}$ of Eq. (1.37) corresponds to elliptic coordinates in the plane. They can be written as

$$
\begin{equation*}
x=D \cosh \xi \cos \eta, \quad y=D \sinh \xi \sin \eta . \tag{1.41}
\end{equation*}
$$

The results are presented in Table 1.
1.3. Separable Coordinate Systems in the Pseudo-Euclidean Plane. The Lie algebra of the isometry group $E(1,1)$ can be represented by

$$
\begin{equation*}
K=\left(t \partial_{x}+x \partial_{t}\right), \quad P_{0}=\partial_{t}, \quad P_{1}=\partial_{x} \tag{1.42}
\end{equation*}
$$

The second-order operator (1.21) is

$$
\begin{align*}
\hat{Q}=a K^{2}+b_{0}\left(K P_{0}+P_{0} K\right)+b_{1}\left(K P_{1}\right. & \left.+P_{1} K\right)+ \\
& +c_{0} P_{0}^{2}+c_{1} P_{1}^{2}+2 c_{2} P_{0} P_{1} .
\end{align*}
$$

Equivalently, the matrix $A$ of (1.24) is

$$
A=\left(\begin{array}{cc}
a & b^{T}  \tag{1.44}\\
b & c
\end{array}\right), \quad C=\left(\begin{array}{cc}
c_{0} & c_{2} \\
c_{2} & c_{1}
\end{array}\right), \quad b^{T}=\left(b_{0}, b_{1}\right) .
$$

We will classify the operators $\hat{Q}$ into conjugate classes and the action of the group $E(1,1)$, including the reflactions

$$
\begin{equation*}
\Pi_{0}:(x, t) \rightarrow(x,-t), \quad \Pi_{1}:(x, t) \rightarrow(-x, t) . \tag{1.45}
\end{equation*}
$$

An element of $E(1,1)$, acting on the Lie algebra $\left(K, P_{0}, P_{1}\right)$ can be represented as

$$
g=\left(\begin{array}{cc}
1 & \xi^{T}  \tag{1.46}\\
0 & \Lambda
\end{array}\right), \quad \xi^{T}=\left(\xi_{0}, \xi_{1}\right), \quad \Lambda \in \mathbb{R}^{2}, \quad \Lambda^{T} J \Lambda=J
$$

with

$$
J=\left(\begin{array}{cc}
1 & 0  \tag{1.47}\\
0 & -1
\end{array}\right)
$$

The matrix A of (1.44) is subject to the transformation (1.26) and in this case we have

$$
A^{\prime}=g^{T} A g=\left(\begin{array}{cc}
a & a \xi^{T}+\beta^{T} \Lambda  \tag{1.48}\\
\xi a+\Lambda^{T} \beta & \Lambda^{T} C \Lambda+\Lambda^{T} \beta \xi^{T}+\xi \beta^{T} \Lambda+a \xi \xi^{T}
\end{array}\right) .
$$

Table 1. Orthogonal coordinate systems on two-dimensional Euclidean plane $E_{2}$

| Coordiante systems | Integrals of motion | Solution of Helmholtz equation |
| :---: | :---: | :---: |
| $\begin{aligned} & \text { I. Cartesian } x, y \\ & -\infty<x, y<\infty \\ & \hline \end{aligned}$ | $Q_{C}=P_{1}^{2}-P_{2}^{2}$ | Exponential functions |
| II. Polar $\begin{aligned} & x=r \cos \varphi, \quad y=r \sin \varphi \\ & 0 \leq r<\infty, 0 \leq \varphi \leq 2 \pi \end{aligned}$ | $Q_{R}=L^{2}$ | Product of Bessel function and exponential |
| III. Parabolic $\begin{aligned} & x=\left(u^{2}-v^{2}\right) / 2, \quad y=u v \\ & 0 \leq u<\infty,-\infty<v<\infty \end{aligned}$ | $Q_{P}=L P_{2}+P_{2} L$ | Product of two parabolic cylinder functions |
| IV. Elliptic $\begin{gathered} x=D \cosh \xi \cos \eta, \\ y=D \sinh \xi \sin \eta \\ 0 \leq \xi<\infty, 0 \leq \eta<2 \pi \end{gathered}$ | $Q_{E}=L^{2}-D^{2} P_{2}^{2}$ | Product of periodic and nonperiodic Mathieu functions |

One of the invariants of this transformation is the constant $a$ which can be chosen to be $a=1$, or is already $a=0$.

Let us first consider $a \neq 0$. Choosing $\xi=-\beta^{T} \Lambda$ and putting $a=1$, we obtain

$$
A^{\prime}=\left(\begin{array}{cc}
1 & 0  \tag{1.49}\\
0 & C^{\prime}
\end{array}\right), \quad C^{\prime}=J \Lambda^{-1} J\left(C-\beta \beta^{T}\right) \Lambda
$$

Notice that $C^{\prime}$ and $C$ are symmetric matrices, but we have

$$
\begin{equation*}
X \equiv J\left(C-\beta \beta^{T}\right), \quad J X^{T}=X J \tag{1.50}
\end{equation*}
$$

that is, $X$ is an element of the Jordan algebra $j o(1,1)$. Since $\Lambda$ is an element of the Lie group $O(1,1)$, we are faced with a well-known problem: the classification of elements of a Jordan algebra with respect to conjugation under the corresponding Lie group. The results are known for all classical Lie and Jordan algebras [35], and for $j o(1,1)$ they are quite simple. The matrix $X$ can be transformed into one of the following

$$
\begin{gather*}
X_{1}=\left(\begin{array}{cc}
p & 0 \\
0 & q
\end{array}\right), \quad X_{2}=\left(\begin{array}{cc}
p & q \\
-q & p
\end{array}\right),  \tag{1.51}\\
X_{3}=\left(\begin{array}{cc}
p+\epsilon & \epsilon \\
-\epsilon & p-\epsilon
\end{array}\right), \quad q>0, \quad \epsilon= \pm 1,
\end{gather*}
$$

with $p \in \mathbb{R}, q \in \mathbb{R}$.

For $a=0,\left(b_{0}, b_{1}\right) \neq(0,0)$ transformation (1.26) leads to Eq. (1.48) in which we set $a=0$. We choose the matrix $\Lambda$ contained in $O(1,1)$ to transform $\Lambda^{T} \beta$ to standard form and then choose $\xi$ to simplify the matrix $C$. For $\left|b_{0}\right|>\left|b_{1}\right|$, $\left|b_{0}\right|<\left|b_{1}\right|$ and $\left|b_{0}\right|=\left|b_{1}\right|$ we can transform $A$ into

$$
\begin{align*}
& A_{1}=\left(\begin{array}{ccc}
0 & \epsilon \sqrt{b_{0}^{2}-b_{1}^{2}} & 0 \\
\epsilon \sqrt{b_{0}^{2}-b_{1}^{2}} & 0 & 0 \\
0 & 0 & 0
\end{array}\right), \quad \epsilon= \pm 1,  \tag{1.52}\\
& A_{2}=\left(\begin{array}{ccc}
0 & 0 & \epsilon \sqrt{b_{1}^{2}-b_{0}^{2}} \\
0 & 0 & 0 \\
\epsilon \sqrt{b_{1}^{2}-b_{0}^{2}} & 0 & 0
\end{array}\right), \quad \epsilon= \pm 1,  \tag{1.53}\\
& A_{3}=\left(\begin{array}{ccc}
0 & 1 & 1 \\
1 & \gamma & -\gamma \\
1 & -\gamma & \gamma
\end{array}\right), \quad \gamma=0,1, \tag{1.54}
\end{align*}
$$

respectively.
Finally, for $a=b_{0}=b_{1}=0, C \neq 0$ we can use $\Lambda$ to transform $C$ into one of its standard forms $J X_{i}, i=1,2,3$ with $X_{i}$ as in (1.51).

Thus, we have obtained a classification of matrices $A$ that determine the operator $\hat{Q}$. Let us now list the corresponding operators. We first notice that $a=b_{1}=b_{0}=0$. The corresponding operator $\hat{Q}$ is in the enveloping algebra of a maximal Abelian subalgebra of $e(1,1)$, namely $\left(P_{0}, P_{1}\right)$. Similarly, for $a=1$ and $X=X_{1}$ in Eq. (1.51) with $p=q$ we find that $Q=K^{2}$ is in the enveloping algebra of a different maximal Abelian subalgebra of $e(1,1)$, namely $o(1,1)$ (generated by $K$ ). These two cases correspond to subgroup type coordinates, the other ones, to nonsubgroup type.

The list of operators must be further simplifield by linear combinations with the Casimir operator

$$
\begin{equation*}
C=P_{0}^{2}-P_{1}^{2} \tag{1.55}
\end{equation*}
$$

Finally, we obtain a representative list of 11 second-order operators in the enveloping algebra of the Lie algebra $e(1,1)$.

$$
\begin{aligned}
Q_{1}(a, b)= & a\left(P_{0}^{2}+P_{1}^{2}\right)+2 b P_{0} P_{1} \\
& (a, b)=(1,0),(1,1), \text { or }(0,1) \\
Q_{2}= & K^{2} \\
Q_{3}= & K P_{1}+P_{1} K \\
Q_{4}= & K P_{0}+P_{0} K
\end{aligned}
$$

$$
\begin{align*}
Q_{5} & =K\left(P_{0}+P_{1}\right)+\left(P_{0}+P_{1}\right) K, \\
Q_{6} & =K\left(P_{0}+P_{1}\right)+\left(P_{0}+P_{1}\right) K+\left(P_{0}-P_{1}\right)^{2},  \tag{1.56}\\
Q_{7} & =K^{2}-l^{2} P_{0} P_{1}, \quad l>0, \\
Q_{8} & =K^{2}-D^{2} P_{1}^{2}, \quad D>0, \\
Q_{9} & =K^{2}-d^{2} P_{1}^{2}, \quad d>0, \\
Q_{10} & =K^{2}+\left(P_{0}+P_{1}\right)^{2}, \\
Q_{11} & =K^{2}-\left(P_{0}+P_{1}\right)^{2} .
\end{align*}
$$

To obtain separable coordinates we proceed as in Sec.1.1.

1. The operator $Q_{1}(a, b)$ for any $a$ and $b$, corresponds to Cartesian coordinates $(t, x)$, since the operators that are really diagonalized are $P_{0}$ and $P_{1}$ (they correspond to a maximal Abelian subalgebra $\left\{P_{0}, P_{1}\right\} \in e(1,1)$ ).
2. The operator $Q_{2}=K^{2}$ also corresponds to subgroup type coordinates, namely pseudopolar coordinates

$$
\begin{array}{lll}
t=r \cosh \alpha, & x=r \sinh \alpha,  \tag{1.57}\\
0 \leq r<\infty, & \infty<\alpha<\infty .
\end{array}
$$

These coordinates only cover part of the pseudo-Euclidean plane, since we have $t^{2}-x^{2}=r^{2} \geq 0$. By interchanging $t$ and $x$ in (1.57) we can parametrize the part with $t^{2}-x^{2}=-r^{2}$.

The operators $Q_{3}, \ldots Q_{11}$ can lead to separable coordinates via the algorithm of Eq.(1.17). Two problems can and do occur. The first is that the roots of Eq. (1.17) may coincide: $\rho_{1}=\rho_{2}$. Then we do not obtain separable coordiantes. This happens in precisely one case, namely that of the operator $Q_{5}$.

To other problem that may occur is that the eigenvalues $\rho(t, x)$ may be complex at least in a part of the $(x, t)$ plane. This part of the plane will then not be covered by the corresponding coordinates $\left(\rho_{1}, \rho_{2}\right)$.

The results of this analysis are presented in Table 2 and essentially agree with those of Kalnins [36].
1.4. The Systems of Coordinates on $S_{2}$. The Lie algebra of isometry group $O(3)$ is given by

$$
\begin{equation*}
L_{i}=-\epsilon_{i k j} u_{k} \frac{\partial}{\partial u_{j}}, \quad\left[L_{i}, L_{k}\right]=\epsilon_{i k j} L_{j}, \quad i, k, j=1,2,3, \tag{1.58}
\end{equation*}
$$

where $u_{i}$ are the Cartesian coordinates in the ambient Euclidean space $E_{3}$. On the sphere $S_{2}$ we have $u_{1}^{2}+u_{2}^{2}+u_{3}^{2}=R^{2}$. The Casimir operator is

$$
\begin{equation*}
C=R^{2} \Delta_{\mathrm{LB}}=L_{1}^{2}+L_{2}^{2}+L_{3}^{2} \tag{1.59}
\end{equation*}
$$

Table 2. Orthogonal coordinate systems on the two-dimensional pseudo-Euclidean plane $E_{1,1}$

| Coordinate system | Integrals of motion | Solution of Helmholtz equation |
| :---: | :---: | :---: |
| I. Cartesian $t, x$ | $Q_{C}=P_{0} P_{1}$ | Product of exponentials |
| $\begin{aligned} & \text { II. Pseudo-polar } \\ & t=r \cosh \tau_{2}, x=r \sinh \tau_{2} \\ & r \geq 0,-\infty<\tau_{2}<\infty \\ & \hline \end{aligned}$ | $Q_{S}=K^{2}$ | Product of Bessel function and exponential |
| $\begin{aligned} & \text { III. Parabolic of type I } \\ & t=1 / 2\left(u^{2}+v^{2}\right), x=u v \\ & v \geq 0,-\infty<u<\infty \\ & \hline \end{aligned}$ | $Q_{P}^{\mathrm{I}}=\left\{P_{1}, K\right\}$ | Product of parabolic cylinder function for $t^{2}-x^{2}>0$ |
| $\begin{aligned} & \text { IV. Parabolic of type II } \\ & t=u v, \quad x=1 / 2\left(u^{2}+v^{2}\right) \\ & v \geq 0,-\infty<u<\infty \\ & \hline \end{aligned}$ | $Q_{P}^{\mathrm{II}}=\left\{P_{0}, K\right\}$ | Product of parabolic cylinder function for $x^{2}-t^{2}>0$ |
| $\begin{aligned} & \text { V. Parabolic of type III } \\ & t=1 / 2(\eta-\zeta)^{2}-(\eta+\zeta), \\ & x=1 / 2(\eta-\zeta)^{2}+(\eta+\zeta) \\ & -\infty<\eta, \zeta<\infty \end{aligned}$ | $\begin{aligned} & Q_{P}^{\mathrm{III}}=\left\{P_{0}, K\right\}+\left\{P_{1}, K\right\} \\ & +\left(P_{0}-P_{1}\right)^{2} \end{aligned}$ | Products of two linear combinations of Airy functions for $x+t>0$ |
| $\begin{aligned} & \text { VI. Hyperbolic of type I } \\ & t=1 / 2\left(\cosh \frac{\eta-\zeta}{2}+\sinh \frac{\eta+\zeta}{2}\right), \\ & x=1 / 2\left(\cosh \frac{\eta-\zeta}{2}-\sinh \frac{\eta+\zeta}{2}\right) \\ & -\infty<\eta, \zeta<\infty \end{aligned}$ | $Q_{H}^{\mathrm{I}}=K^{2}-l^{2} P_{0} P_{1}$ | Product of Mathieu equation solutions with argument displaced by $i \pi / 2$ |
| $\begin{aligned} & \text { VII. Hyperbolic of type II } \\ & t=\left(\sinh (\eta-\zeta)+\mathrm{e}^{\eta+\zeta}\right), \\ & x=\left(\sinh (\eta-\zeta)-\mathrm{e}^{\eta+\zeta}\right) \\ & -\infty<\eta, \zeta<\infty \end{aligned}$ | $Q_{H}^{\mathrm{II}}=K^{2}+\left(P_{1}+P_{2}\right)^{2}$ | Product of two solutions of Bessel's equation, one with real and one with imaginary arguments |
| $\begin{aligned} & \text { VIII. Hyperbolic of type III } \\ & t=\left(\cosh (\eta-\zeta)+\mathrm{e}^{\eta+\zeta}\right), \\ & x=\left(\cosh (\eta-\zeta)-\mathrm{e}^{\eta+\zeta}\right) \\ & -\infty<\eta, \zeta<\infty \end{aligned}$ | $Q_{H}^{\text {III }}=K^{2}-\left(P_{1}+P_{2}\right)^{2}$ | Product of two solutions of Bessel's equation |
| $\begin{aligned} & \hline \text { IX. Elliptic of type I } \\ & t=D \sinh \eta \cosh \zeta, \\ & x=D \cosh \eta \sinh \zeta \\ & -\infty<\eta, \zeta<\infty \\ & \hline \end{aligned}$ | $Q_{E}^{\mathrm{I}}=K^{2}+D^{2} P_{1}^{2}$ | Product of two solutions of the nonperiodic Mathieu equation |
| $\begin{aligned} & \text { X. Elliptic of type II } \\ & \text { (i) } t=d \cosh \eta \cosh \zeta, \\ & x=d \sinh \eta \sinh \zeta \\ & -\infty<\eta<\infty, \zeta \geq 0 \\ & \text { (ii) } t=d \cos \eta \cos \zeta \\ & x=d \sin \eta \sin \zeta \\ & 0<\eta<2 \pi, 0 \leq \zeta<\pi \\ & \hline \end{aligned}$ | $Q_{E}^{\mathrm{II}}=K^{2}-d^{2} P_{1}^{2}$ | Product of two solutions <br> (i) of the nonperiodic <br> Mathieu equation <br> (ii) of the periodic <br> Mathieu equation |

and the Laplace-Beltrami Eq. (1.10) for $S_{2}$ has the form

$$
\begin{equation*}
\Delta_{\mathrm{LB}} \Psi=-\frac{\ell(\ell+1)}{R^{2}} \Psi, \quad \Psi_{l k}(\alpha, \beta)=\Xi_{l k}(\alpha) \Phi_{l k}(\beta), \tag{1.60}
\end{equation*}
$$

where $\ell=0,1,2, \ldots$ The second-order operator $\hat{Q}$ of Eq. (1.21) is given by

$$
\begin{equation*}
\hat{Q}=A_{i k} L_{i} L_{k}, \quad A_{i k}=A_{k i} \tag{1.61}
\end{equation*}
$$

The transformation matrix for $O(3)$ can be represented as

$$
\begin{align*}
& \quad g= \\
& =\left(\begin{array}{ccc}
\cos \alpha \cos \beta \cos \gamma-\sin \alpha \sin \gamma & -\cos \alpha \cos \beta \sin \gamma-\sin \alpha \cos \gamma & \cos \alpha \sin \beta \\
\sin \alpha \cos \beta \cos \gamma+\cos \alpha \sin \gamma & -\sin \alpha \cos \beta \sin \gamma+\cos \alpha \cos \gamma & \sin \alpha \sin \beta \\
-\sin \beta \cos \gamma & \sin \beta \sin \gamma & \cos \beta
\end{array}\right), \tag{1.62}
\end{align*}
$$

where $(\alpha, \beta, \gamma)$ are the Euler angles.
The matrix $A_{i k}$ can be diagonalized to give

$$
\begin{equation*}
\hat{Q}\left(a_{1}, a_{2}, a_{3}\right) \equiv Q=a_{1} L_{1}^{2}+a_{2} L_{2}^{2}+a_{3} L_{3}^{2} \tag{1.63}
\end{equation*}
$$

For $a_{1}=a_{2}=a_{3}$ we have $Q \sim 0$. If two eigenvalues of $A_{i k}$ are equal, e.g., $a_{1}=a_{2} \neq a_{3}$, or $a_{1} \neq a_{2}=a_{3}$, or $a_{1}=a_{3} \neq a_{2}$, we can transform the operator $Q$ into the operators: $Q(0,0,1)=L_{3}^{2}, Q(1,0,0)=L_{1}^{2}$ or $Q(0,1,0)=L_{2}^{2}$, respectively. The corresponding separable coordinates on $S_{2}$ are the three types of spherical ones

$$
\begin{align*}
& u_{1}=R \sin \theta \cos \varphi=R \cos \theta^{\prime}=R \sin \theta^{\prime \prime} \sin \varphi^{\prime \prime} \\
& u_{2}=R \sin \theta \sin \varphi=R \sin \theta^{\prime} \cos \varphi^{\prime}=R \cos \theta^{\prime \prime}  \tag{1.64}\\
& u_{3}=R \cos \theta=R \sin \theta^{\prime} \sin \varphi^{\prime}=R \sin \theta^{\prime \prime} \cos \varphi^{\prime \prime}
\end{align*}
$$

where $\varphi \in[0,2 \pi), \theta \in[0, \pi]$. They correspond to the group reduction $O(3) \supset$ $O(2)$ and $X=L_{i}^{2}$ is invariant under $O(2)$ and under reflections in all coordinate planes.

The $O(3)$ unitary irreducible representation matrix elements of (1.62) result in the well-known transformation formula for spherical functions $Y_{l m}(\theta, \varphi)[8,37]$

$$
\begin{equation*}
Y_{l, m^{\prime}}\left(\theta^{\prime}, \varphi^{\prime}\right)=\sum_{m=-l}^{l} D_{m m^{\prime}}^{l}(\alpha, \beta \gamma) Y_{l, m}(\theta, \varphi) \tag{1.65}
\end{equation*}
$$

where $D_{m_{1}, m_{2}}^{l}(\alpha, \beta, \gamma)$ are the Wigner $D$-functions

$$
\begin{equation*}
D_{m, m^{\prime}}^{\ell}(\alpha, \beta, \gamma)=\mathrm{e}^{-i m \alpha} d_{m, m^{\prime}}^{\ell}(\beta) \mathrm{e}^{-i m^{\prime} \gamma} \tag{1.66}
\end{equation*}
$$

$$
\begin{align*}
d_{m, m^{\prime}}^{\ell}(\beta)= & \frac{(-1)^{m-m^{\prime}}}{\left(m-m^{\prime}\right)!} \sqrt{\frac{(\ell+m)!\left(\ell-m^{\prime}\right)!}{(\ell-m)!\left(\ell+m^{\prime}\right)!}}\left(\cos \frac{1}{2} \beta\right)^{2 \ell-m+m^{\prime}} \times \\
& \times\left(\sin \frac{1}{2} \beta\right)^{m-m^{\prime}} F\left[\begin{array}{c}
m-\ell,-m^{\prime}-\ell \\
m-m^{\prime}+1
\end{array} ;-\tan ^{2} \frac{1}{2} \beta\right] \tag{1.67}
\end{align*}
$$

and the spherical angles $(\theta, \varphi)$ and $\left(\theta^{\prime}, \varphi^{\prime}\right)$ are related by

$$
\begin{align*}
& \cos \theta^{\prime}=\cos \theta \cos \beta+\sin \theta \sin \beta \cos (\varphi-\alpha) \\
& \cot \left(\varphi^{\prime}+\gamma\right)=\cot (\varphi-\alpha) \cos \beta-\frac{\cot \theta \sin \beta}{\sin (\varphi-\alpha)} \tag{1.68}
\end{align*}
$$

In particular, $Y_{l m}(\theta, \varphi)$ corresponding to the solution of Laplace-Beltrami equation in the systems of coordinates (1.64) are related by the formulas

$$
\begin{gather*}
Y_{l, m^{\prime}}\left(\theta^{\prime}, \varphi^{\prime}\right)=\sum_{m=-l}^{l} D_{m m^{\prime}}^{l}\left(0, \frac{\pi}{2}, \frac{\pi}{2}\right) Y_{l, m}(\theta, \varphi),  \tag{1.69}\\
Y_{l, m^{\prime \prime}}\left(\theta^{\prime \prime}, \varphi^{\prime \prime}\right)=\sum_{m=-l}^{l} D_{m m^{\prime \prime}}^{l}\left(\frac{\pi}{2}, \frac{\pi}{2}, 0\right) Y_{l, m}(\theta, \varphi),  \tag{1.70}\\
Y_{l, m^{\prime \prime}}\left(\theta^{\prime \prime}, \varphi^{\prime \prime}\right)=\sum_{m^{\prime}=-l}^{l} D_{m^{\prime} m^{\prime \prime}}^{l}\left(0, \frac{\pi}{2}, \frac{\pi}{2}\right) Y_{l, m^{\prime}}\left(\theta^{\prime}, \varphi^{\prime}\right) . \tag{1.71}
\end{gather*}
$$

When all three eigenvalues $a_{i}$ are different, then the separable coordinates in Eq. (1.60) are elliptic ones [38-40]. These can be written in algebraic form, as

$$
\begin{gather*}
u_{1}^{2}=R^{2} \frac{\left(\rho_{1}-a_{1}\right)\left(\rho_{2}-a_{1}\right)}{\left(a_{2}-a_{1}\right)\left(a_{3}-a_{1}\right)}, \quad u_{2}^{2}=R^{2} \frac{\left(\rho_{1}-a_{2}\right)\left(\rho_{2}-a_{2}\right)}{\left(a_{3}-a_{2}\right)\left(a_{1}-a_{2}\right)}  \tag{1.72}\\
u_{3}^{2}=R^{2} \frac{\left(\rho_{1}-a_{3}\right)\left(\rho_{2}-a_{3}\right)}{\left(a_{1}-a_{3}\right)\left(a_{2}-a_{3}\right)}
\end{gather*}
$$

with $a_{1} \leq \rho_{1} \leq a_{2} \leq \rho_{2} \leq a_{3}$.
In trigonometric form we put

$$
\begin{equation*}
\rho_{1}=a_{1}+\left(a_{2}-a_{1}\right) \cos ^{2} \phi, \quad \rho_{2}=a_{3}-\left(a_{3}-a_{2}\right) \cos ^{2} \theta \tag{1.73}
\end{equation*}
$$

and obtain

$$
\begin{gather*}
u_{1}=R \sqrt{1-k^{2} \cos ^{2} \theta} \cos \phi, \quad u_{2}=R \sin \theta \sin \phi  \tag{1.74}\\
u_{3}=R \sqrt{1-k^{2} \cos ^{2} \phi} \cos \theta, \quad 0 \leq \phi<2 \pi, \quad 0 \leq \theta \leq \pi
\end{gather*}
$$

Table 3. Orthogonal coordinate systems on two-dimensional sphere $S_{2}$

| Coordinate systems | Integrals of motion | Solution of Helmholtz equation | Limiting systems on $E_{2}$ |
| :---: | :---: | :---: | :---: |
| I. Spherical $\begin{aligned} & u_{1}=R \sin \theta \cos \varphi, \\ & u_{2}=R \sin \theta \sin \varphi, \\ & u_{3}=R \cos \theta \\ & \varphi \in[0,2 \pi), \theta \in[0, \pi] \end{aligned}$ | $Q_{S}=L_{3}^{2}$ | Product of associated Legendre polynomials and exponential | Polar <br> Cartesian |
| II. Elliptic $\begin{aligned} & u_{1}=R \operatorname{sn}(\alpha, k) \operatorname{dn}\left(\beta, k^{\prime}\right), \\ & u_{2}=R \operatorname{cn}(\alpha, k) \operatorname{cn}\left(\beta, k^{\prime}\right), \\ & u_{3}=R \operatorname{dn}(\alpha, k) \operatorname{sn}\left(\beta, k^{\prime}\right) \\ & \alpha \in[-K, K], \\ & \beta \in\left[-2 K^{\prime}, 2 K^{\prime}\right] \end{aligned}$ <br> *After rotation. | $Q_{E}=k^{2} L_{3}^{2}-k^{2} L_{1}^{2}$ | Product of two <br> Lamé polynomials | Elliptic <br> Polar <br> Cartesian <br> Parabolic* |

where

$$
\begin{equation*}
k^{2}=\frac{a_{2}-a_{1}}{a_{3}-a_{1}}=\sin ^{2} f, \quad k^{\prime 2}=\frac{a_{3}-a_{2}}{a_{3}-a_{1}}=\cos ^{2} f, \quad k^{2}+k^{\prime 2}=1 . \tag{1.75}
\end{equation*}
$$

The Jacobi elliptic version of elliptic coordinates is obtained by putting

$$
\begin{equation*}
\rho_{1}=a_{1}+\left(a_{2}-a_{1}\right) \operatorname{sn}^{2}(\alpha, k), \quad \rho_{2}=a_{2}+\left(a_{3}-a_{2}\right) \mathrm{cn}^{2}\left(\beta, k^{\prime}\right) . \tag{1.76}
\end{equation*}
$$

We obtain

$$
\begin{gather*}
u_{1}=R \operatorname{sn}(\alpha, k) \operatorname{dn}\left(\beta, k^{\prime}\right), \quad u_{2}=R \operatorname{cn}(\alpha, k) \operatorname{cn}\left(\beta, k^{\prime}\right)  \tag{1.77}\\
u_{3}=R \operatorname{dn}(\alpha, k) \operatorname{sn}\left(\beta, k^{\prime}\right), \quad-K \leq \alpha \leq K, \quad-2 K^{\prime} \leq \beta \leq 2 K^{\prime}
\end{gather*}
$$

where $\operatorname{sn}(\alpha, k)$, $\mathrm{cn}\left(\alpha, k^{\prime}\right)$ and $\operatorname{dn}(\beta, k)$ are the Jacobi elliptic functions with modulus $k$ and $k^{\prime}$, and $K$ and $K^{\prime}$ are the complete elliptic integrals [41].

The interfocal distance for the ellipses on the upper hemisphere is equal to $2 f R$. The results are given in Table 3 .
1.5. Systems of Coordinates on $H_{2}$. The isometry group for the hyperboloid $H_{2}: u_{0}^{2}-u_{1}^{2}-u_{2}^{2}=R^{2}$, where $u_{i}(i=0,1,2)$ are the Cartesian coordinates in the ambient space $E_{2,1}$, is $O(2,1)$. We choose a standard basis $K_{1}, K_{2}, L_{3}$ for the Lie algebra $o(2,1)$ :

$$
K_{1}=-\left(u_{0} \partial_{u_{2}}+u_{2} \partial_{u_{0}}\right), K_{2}=-\left(u_{0} \partial_{u_{1}}+u_{1} \partial_{u_{0}}\right), L_{3}=-\left(u_{1} \partial_{u_{2}}-u_{2} \partial_{u_{1}}\right)
$$

with commutation relations

$$
\begin{equation*}
\left[K_{1}, K_{2}\right]=-L_{3}, \quad\left[L_{3}, K_{1}\right]=K_{2}, \quad\left[K_{2}, L_{3}\right]=K_{1} . \tag{1.78}
\end{equation*}
$$

The Casimir operator is

$$
\begin{equation*}
C=R^{2} \Delta_{\mathrm{LB}}=K_{1}^{2}+K_{2}^{2}-L_{3}^{2} \tag{1.79}
\end{equation*}
$$

and the Laplace-Beltrami equation (1.10) is given by

$$
\begin{equation*}
\Delta_{\mathrm{LB}} \Psi=\frac{l(l+1)}{R^{2}} \Psi, \quad \Psi_{l \lambda}\left(\zeta_{1}, \zeta_{2}\right)=\Xi_{l \lambda}\left(\zeta_{1}\right) \Phi_{l \lambda}\left(\zeta_{2}\right) \tag{1.80}
\end{equation*}
$$

where $\ell$ for principal series of the unitary irreducible representations has the form

$$
\begin{equation*}
\ell=-\frac{1}{2}+i \rho, \quad 0<\rho<\infty \tag{1.81}
\end{equation*}
$$

The second-order operator $\hat{Q}$ of Eq. (1.21):

$$
\begin{align*}
Q=a K_{1}^{2}+b\left(K_{1} K_{2}+K_{2} K_{1}\right)+c K_{2}^{2}+ & d\left(K_{1} L_{3}+L_{3} K_{1}\right)+ \\
& +e\left(K_{2} L_{3}+L_{3} K_{2}\right)+f L_{3}^{2} \tag{1.82}
\end{align*}
$$

can be used to classify all coordinate systems on $H_{2}$. The classification of the operators $Q$ can be reduced to a classification of the normal forms of the elements of the Jordan algebra $j o(2,1)$ [35]. There are 9 inequivalent forms, in one-to-one correspondence with the 9 existing separable coordinate systems [15, 42, 43]. All the coordinate systems are orthogonal ones.

The normal forms of the operator $Q$ and the corresponding coordinates are given in Table 4. Cases I, II, and III correspond to subgroup type coordinates. The corresponding subgroups are $O(2), O(1,1)$ and $E(1)$, respectively. The $O(1,1)$ subgroup in the equidistant coordinates acts in the 01 plane. We could also have chosen the 02 plane (i.e., permuted $u_{1}$ and $u_{2}$ ).

The elliptic and hyperbolic coordinates of cases IV and V are given in algebraic form. Equivalently, they can be expressed, e.g., in terms of Jacobi elliptic functions. This makes it possible to express the coordinates in the ambient space directly, rather than their squares. Indeed, if we put

$$
\begin{equation*}
\varrho_{1}=a_{1}-\left(a_{1}-a_{3}\right) \operatorname{dn}^{2}(\alpha, k), \quad \varrho_{2}=a_{1}-\left(a_{1}-a_{2}\right) \operatorname{sn}^{2}\left(\beta, k^{\prime}\right) \tag{1.83}
\end{equation*}
$$

and

$$
\begin{equation*}
k^{2}=\frac{a_{2}-a_{3}}{a_{1}-a_{3}}, \quad k^{\prime 2}=\frac{a_{1}-a_{2}}{a_{1}-a_{3}}, \quad k^{2}+k^{\prime 2}=1 \tag{1.84}
\end{equation*}
$$

into the expressions

$$
\begin{gather*}
u_{0}^{2}=R^{2} \frac{\left(\varrho_{1}-a_{3}\right)\left(\varrho_{2}-a_{3}\right)}{\left(a_{1}-a_{3}\right)\left(a_{2}-a_{3}\right)}, \quad u_{1}^{2}=R^{2} \frac{\left(\varrho_{1}-a_{2}\right)\left(\varrho_{2}-a_{2}\right)}{\left(a_{2}-a_{3}\right)\left(a_{1}-a_{2}\right)}  \tag{1.85}\\
u_{2}^{2}=R^{2} \frac{\left(\varrho_{1}-a_{1}\right)\left(a_{1}-\varrho_{2}\right)}{\left(a_{1}-a_{2}\right)\left(a_{1}-a_{3}\right)}
\end{gather*}
$$

Table 4. Orthogonal systems of coordinate on two-dimensional hyperboloid $\mathrm{H}_{2}$

| Coordinate systems and integrals of motion | Coordinate | Limiting systems on $E_{2}$ | Limiting systems on $E_{1,1}$ |
| :---: | :---: | :---: | :---: |
| I. Pseudo-spherical $\begin{aligned} & \tau>0, \varphi \in[0,2 \pi) \\ & Q_{S}=L_{3}^{2} \end{aligned}$ | $\begin{aligned} & u_{0}=R \cosh \tau \\ & u_{1}=R \sinh \tau \cos \varphi \\ & u_{2}=R \sinh \tau \sin \varphi \end{aligned}$ | Polar | Cartesian |
| $\begin{aligned} & \text { II. Equidistant } \\ & \tau_{1,2} \in \mathbb{R} \\ & Q_{E Q}=K_{1}^{2} \end{aligned}$ | $\begin{aligned} & u_{0}=R \cosh \tau_{1} \cosh \tau_{2} \\ & u_{1}=R \cosh \tau_{1} \sinh \tau_{2} \\ & u_{2}=R \sinh \tau_{1} \end{aligned}$ | Cartesian | Polar |
| III. Horocyclic $\begin{aligned} & \tilde{y}>0, \tilde{x} \in \mathbb{R} \\ & Q_{H O}=\left(K_{1}+L_{3}\right)^{2} \end{aligned}$ | $\begin{aligned} & u_{0}=R\left(\tilde{x}^{2}+\tilde{y}^{2}+1\right) / 2 \tilde{y} \\ & u_{1}=R\left(\tilde{x}^{2}+\tilde{y}^{2}-1\right) / 2 \tilde{y} \\ & u_{2}=R \tilde{x} / \tilde{y} \end{aligned}$ | Cartesian | Rectangular Cartesian coordinates rotated by $\pi / 4$ (nonorthogonal) |
| $\begin{aligned} & \text { IV. Elliptic } \\ & a_{3}<a_{2}<\rho_{2}<a_{1}<\rho_{1} \\ & Q_{E}=L_{3}^{2}+\sinh ^{2} f K_{2}^{2} \end{aligned}$ | $\begin{aligned} & u_{0}^{2}=R^{2} \frac{\left(\rho_{1}-a_{3}\right)\left(\rho_{2}-a_{3}\right)}{\left(a_{1}-a_{3}\right)\left(a_{2}-a_{3}\right)} \\ & u_{1}^{2}=R^{2} \frac{\left(\rho_{1}-a_{2}\right)\left(\rho_{2}-a_{2}\right)}{\left(a_{1}-a_{2}\right)\left(a_{2}-a_{3}\right)} \\ & u_{2}^{2}=R^{2} \frac{\left(\rho_{1}-a_{1}\right)\left(a_{1}-\rho_{2}\right)}{\left(a_{1}-a_{2}\right)\left(a_{1}-a_{3}\right)} \end{aligned}$ | Elliptic <br> Parabolic <br> Cartesian | Elliptic I, II, III <br> Cartesian |
| $\begin{aligned} & \text { V. Hyperbolic } \\ & \rho_{2}<a_{3}<a_{2}<a_{1}<\rho_{1} \\ & Q_{H}=K_{2}^{2}-\sin ^{2} \alpha L_{3}^{2} \end{aligned}$ | $\begin{aligned} & u_{0}^{2}=R^{2} \frac{\left(\rho_{1}-a_{2}\right)\left(a_{2}-\rho_{2}\right)}{\left(a_{1}-a_{2}\right)\left(a_{2}-a_{3}\right)} \\ & u_{1}^{2}=R^{2} \frac{\left(\rho_{1}-a_{3}\right)\left(a_{3}-\rho_{2}\right)}{\left(a_{1}-a_{3}\right)\left(a_{2}-a_{3}\right)} \\ & u_{2}^{2}=R^{2} \frac{\left(\rho_{1}-a_{1}\right)\left(a_{1}-\rho_{2}\right)}{\left(a_{1}-a_{2}\right)\left(a_{1}-a_{3}\right)} \end{aligned}$ | Cartesian | Elliptic II <br> Parabolic I |
| VI. Semihyperbolic $\begin{aligned} & \mu_{1,2}>0 \\ & Q_{S H}=-\left\{K_{1}, L_{3}\right\} \end{aligned}$ | $\begin{aligned} & \frac{u_{0}^{2}+u_{1}^{2}}{R^{2}}=\left(1+\mu_{1}^{2}\right)\left(1+\mu_{2}^{2}\right) \\ & \frac{u_{0}^{2}-u_{1}^{2}}{R^{2}}=\left(1+\mu_{1} \mu_{2}\right) \\ & u_{2}=R \sqrt{\mu_{1} \mu_{2}} \end{aligned}$ | Parabolic <br> Cartesian | Cartesian |
| VII. Elliptic-parabolic $\begin{aligned} & a \in \mathbb{R}, \vartheta \in(-\pi / 2, \pi / 2) \\ & Q_{E P}=\left(K_{1}+L_{3}\right)^{2}+K_{2}^{2} \end{aligned}$ | $\begin{aligned} & u_{0}=R \frac{\cosh ^{2} a+\cos ^{2} \vartheta}{2 \cosh a \cos \vartheta} \\ & u_{1}=R \frac{\sinh ^{2} a-\sin ^{2} \vartheta}{2 \cosh a \cos \vartheta} \\ & u_{2}=R \tan \vartheta \tanh a \end{aligned}$ | Parabolic | Hyperbolic II |
| VIII. Hyperbolic-parabolic $\begin{aligned} & b>0, \vartheta \in(0, \pi) \\ & Q_{H P}=\left(K_{1}+L_{3}\right)^{2}-K_{2}^{2} \end{aligned}$ | $\begin{aligned} & u_{0}=R \frac{\cosh ^{2} b+\cos ^{2} \vartheta}{2 \sinh b \sin \vartheta} \\ & u_{1}=R \frac{\sinh ^{2} b-\sin ^{2} \vartheta}{2 \sinh b \sin \vartheta} \\ & u_{2}=R \cot \vartheta \operatorname{coth} b \end{aligned}$ | Cartesian | Hyperbolic III |


| Coordinate systems <br> and integrals of motion | Coordinate | Limiting <br> systems on $E_{2}$ | Limiting <br> systems on $E_{1,1}$ |
| :--- | :---: | :--- | :--- |
| IX. Semicircular-parabolic <br> $\xi, \eta>0$ | $u_{0}=R \frac{\left(\xi^{2}+\eta^{2}\right)^{2}+4}{8 \xi \eta}$ | Cartesian | Does not <br> lorrespond to a <br> separable |
| $Q_{S C P}=\left\{K_{1}, K_{2}\right\}+\left\{K_{2}, L_{3}\right\}$ | $u_{1}=R \frac{\left(\xi^{2}+\eta^{2}\right)^{2}-4}{8 \xi \eta}$ |  |  |
| coordinate system |  |  |  |
|  | $u_{2}=R \frac{\eta^{2}-\xi^{2}}{2 \xi \eta}$ |  |  |

we obtain the elliptic coordinates in Jacobi form

$$
\begin{gather*}
u_{0}=R \operatorname{sn}(\alpha, k) \operatorname{dn}\left(\beta, k^{\prime}\right), \quad u_{1}=i R \operatorname{cn}(\alpha, k) \operatorname{cn}\left(\beta, k^{\prime}\right) \\
u_{2}=i R \operatorname{dn}(\alpha, k) \operatorname{sn}\left(\beta, k^{\prime}\right), \quad \alpha \in\left(i K^{\prime}, i K^{\prime \prime}+2 K\right), \quad \beta \in\left[0,4 K^{\prime}\right) . \tag{1.86}
\end{gather*}
$$

## 2. CONTRACTIONS OF THE LIE ALGEBRA AND CASIMIR OPERATOR

2.1. Contractions from $o(3)$ to $e(2)$. We shall use $R^{-1}$ as the contraction parameter. To realize the contraction explicitly, let us introduce homogeneous or Beltrami coordinates on the sphere, putting

$$
\begin{equation*}
x_{\mu}=R \frac{u_{\mu}}{u_{3}}=\frac{u_{\mu}}{\sqrt{1-\left(u_{1}^{2}+u_{2}^{2}\right) / R^{2}}}, \quad \mu=1,2 . \tag{2.1}
\end{equation*}
$$

Geometrically ( $x_{1}, x_{2}$ ) correspond to a projection from the centre of the sphere to a tangent plane at the North pole. In this parametrization the metric tensor has the following form

$$
\begin{equation*}
g_{\mu \nu}=\frac{1}{1+r^{2} / R^{2}}\left[\delta_{\mu \nu}+\frac{x_{\mu} x_{\nu}}{r^{2}} \frac{1}{1+r^{2} / R^{2}}\right], \quad r^{2}=x_{\mu} x_{\mu} \tag{2.2}
\end{equation*}
$$

The Laplace-Beltrami operator corresponds to

$$
\begin{align*}
& \Delta_{\mathrm{LB}}=\left(1+\frac{r^{2}}{R^{2}}\right)\left[\frac{\partial^{2}}{\partial x_{\mu}^{2}}+\frac{x_{\mu}}{R^{2}} \frac{\partial}{\partial x_{\mu}}+\frac{1}{R^{2}}\left(x_{\mu} \frac{\partial}{\partial x_{\mu}}\right)^{2}\right]= \\
&=\left(\pi_{1}^{2}+\pi_{2}^{2}+\frac{L_{3}^{2}}{R^{2}}\right) \tag{2.3}
\end{align*}
$$

where

$$
\begin{equation*}
\pi_{\mu}=\left(\frac{\partial}{\partial x_{\mu}}+\frac{x_{\mu} x_{\nu}}{R^{2}} \frac{\partial}{\partial x_{\nu}}\right), \quad L_{3}=\left(x_{1} \frac{\partial}{\partial x_{2}}-x_{2} \frac{\partial}{\partial x_{1}}\right) . \tag{2.4}
\end{equation*}
$$

Using the connection between operators $\pi_{\mu}$ and the generators of the $O(3)$ group

$$
-\frac{L_{1}}{R}=\pi_{2}, \quad \frac{L_{2}}{R}=\pi_{1}, \quad L_{3}=-\left(x_{1} \pi_{2}-x_{2} \pi_{1}\right)
$$

we obtain the following commutation relations

$$
\begin{equation*}
\left[L_{3}, \pi_{1}\right]=\pi_{2}, \quad\left[L_{3}, \pi_{2}\right]=-\pi_{1}, \quad\left[\pi_{1}, \pi_{2}\right]=\frac{L_{3}}{R^{2}} \tag{2.5}
\end{equation*}
$$

so that for $R \rightarrow \infty$ the $o(3)$ algebra contracts to the $e(2)$ one. Moreover the momenta $\pi_{\mu}$ contract to $P_{\mu}=\partial / \partial x_{\mu},(\mu=1,2)$ and the $o(3)$ Laplace-Beltrami operator (2.3) contracts to the $e(2)$ one:

$$
\begin{equation*}
\Delta_{\mathrm{LB}}=\pi_{1}^{2}+\pi_{2}^{2}+\frac{L_{3}^{2}}{R^{2}} \rightarrow \Delta=\left(P_{1}^{2}+P_{2}^{2}\right) \tag{2.6}
\end{equation*}
$$

2.2. Contractions from $o(2,1)$ to $e(2)$. As in Sec. 2.1, let us introduce the Beltrami coordinates on the hyperboloid $H_{2}$ putting

$$
\begin{equation*}
x_{\mu}=R \frac{u_{\mu}}{u_{0}}=R \frac{u_{\mu}}{\sqrt{R^{2}+u_{1}^{2}+u_{2}^{2}}}, \quad \mu=1,2 \tag{2.7}
\end{equation*}
$$

The $O(2,1)$ generators can be expressed as:

$$
\begin{gathered}
-\frac{K_{1}}{R} \equiv \tilde{\pi}_{2}=\frac{\partial}{\partial x_{2}}-\frac{x_{2}}{R^{2}}\left(x_{1} \frac{\partial}{\partial x_{1}}+x_{2} \frac{\partial}{\partial x_{2}}\right) \\
-\frac{K_{2}}{R} \equiv \tilde{\pi}_{1}=p_{1}-\frac{x_{1}}{R^{2}}\left(x_{1} \frac{\partial}{\partial x_{1}}+x_{2} \frac{\partial}{\partial x_{2}}\right) \\
L_{3}=x_{1} \pi_{2}-x_{2} \pi_{1}
\end{gathered}
$$

The commutation relations of the $o(2,1)$ algebra (1.78) in terms of the new operators take the form

$$
\begin{equation*}
\left[\tilde{\pi}_{1}, \tilde{\pi}_{2}\right]=-\frac{L_{3}}{R^{2}}, \quad\left[L_{3}, \tilde{\pi}_{1}\right]=\tilde{\pi}_{2}, \quad\left[\tilde{\pi}_{2}, L_{3}\right]=\tilde{\pi}_{1} \tag{2.8}
\end{equation*}
$$

so, that for $R \rightarrow \infty$ the $o(2,1)$ algebra contracts to $e(2)$ and the momenta $\tilde{\pi}_{\mu}$ to $P_{\mu}=\partial / \partial x_{\mu}$. The $o(2,1)$ Laplace-Beltrami operator (1.2) contracts to the $e(2)$ one:

$$
\begin{equation*}
\Delta_{\mathrm{LB}}=\tilde{\pi}_{1}^{2}+\tilde{\pi}_{2}^{2}-\frac{L_{3}^{2}}{R^{2}} \rightarrow \Delta=\left(P_{1}^{2}+P_{2}^{2}\right) \tag{2.9}
\end{equation*}
$$

2.3. Contractions from $o(2,1)$ to $e(1,1)$. Let us introduce Beltrami coordinates on hyperboloid $\mathrm{H}_{2}$

$$
\begin{equation*}
y_{0}=R \frac{u_{0}}{u_{2}}=R \frac{u_{0}}{\sqrt{u_{0}^{2}+u_{1}^{2}-R^{2}}}, \quad y_{1}=R \frac{u_{1}}{u_{2}}=R \frac{u_{1}}{\sqrt{u_{0}^{2}+u_{1}^{2}-R^{2}}} \tag{2.10}
\end{equation*}
$$

The $O(2,1)$ generators can be expressed as

$$
\begin{gather*}
-\frac{K_{1}}{R} \equiv \tilde{\tilde{\pi}}_{1}=\frac{\partial}{\partial y_{0}}-\frac{y_{0}}{R^{2}}\left(y_{0} \frac{\partial}{\partial y_{0}}+y_{1} \frac{\partial}{\partial y_{1}}\right), \\
-\frac{L_{3}}{R} \equiv \tilde{\tilde{\pi}}_{2}=\frac{\partial}{\partial y_{1}}+\frac{y_{1}}{R^{2}}\left(y_{0} \frac{\partial}{\partial y_{0}}+y_{1} \frac{\partial}{\partial y_{1}}\right),  \tag{2.11}\\
-K_{2} \equiv K=y_{0} \tilde{\tilde{\pi}}_{2}+y_{1} \tilde{\tilde{\pi}}_{1} .
\end{gather*}
$$

The commutators of the $o(2,1)$ algebra (1.78) in the new operators ( $\tilde{\tilde{\pi}}_{1}, \tilde{\tilde{\pi}}_{2}, K$ ) take the form

$$
\begin{equation*}
\left[\tilde{\pi}_{1}, \tilde{\pi}_{2}\right]=\frac{K}{R^{2}}, \quad\left[K, \tilde{\tilde{\pi}}_{1}\right]=-\tilde{\tilde{\pi}}_{2}, \quad\left[\tilde{\tilde{\pi}}_{2}, K\right]=\tilde{\tilde{\pi}}_{1} \tag{2.12}
\end{equation*}
$$

so, that for $R \rightarrow \infty$ the $o(2,1)$ algebra contracts to the $e(1,1)$ one. The $o(2,1)$ Laplace-Beltrami operator contracts to the $e(1,1)$ one:

$$
\begin{equation*}
\Delta_{\mathrm{LB}}=\tilde{\tilde{\pi}}_{1}^{2}-\tilde{\tilde{\pi}}_{2}^{2}+\frac{K^{2}}{R^{2}} \rightarrow \frac{\partial^{2}}{\partial y_{0}^{2}}-\frac{\partial^{2}}{\partial y_{1}^{2}} \tag{2.13}
\end{equation*}
$$

and Eq. (1.80) transforms for large $\ell \sim p R$ to the one-dimensional Klein-Gordan equation.

$$
\begin{equation*}
\frac{\partial^{2} \psi}{\partial y_{0}^{2}}-\frac{\partial^{2} \psi}{\partial y_{1}^{2}}+p^{2} \psi=0 \tag{2.14}
\end{equation*}
$$

## 3. CONTRACTION FOR SYSTEMS OF COORDINATES

3.1. Contractions and Coordinate Systems on $S_{2}$

1. Spherical coordinates on $S_{2}$ to polar on $E_{2}$. We consider the spherical coordinate (1.64) with the parametr $a_{1}=a_{2}$ and put

$$
\tan \theta=\frac{r}{R}
$$

In the contraction limit $R \rightarrow \infty, \theta \rightarrow 0$ we have

$$
x_{1}=R \frac{u_{1}}{u_{3}} \rightarrow x=r \cos \varphi, \quad x_{2}=R \frac{u_{2}}{u_{3}} \rightarrow y=r \sin \varphi
$$

and

$$
Q_{S}=L_{3}^{2} \rightarrow L^{2}
$$

2. Spherical coordinate on $S_{2}$ to Cartesian on $E_{2}$. We choose the spherical coordinate (1.64) with $a_{2}=a_{3}$. Putting

$$
\cos \theta^{\prime} \sim \frac{x}{R} \sim 0, \quad \cos \varphi^{\prime} \sim \frac{y}{R} \sim 0
$$

and taking the limit $R \rightarrow \infty$ and $\theta^{\prime} \rightarrow \frac{\pi}{2}, \varphi^{\prime} \rightarrow \frac{\pi}{2}$, we obtain

$$
\frac{1}{R^{2}} Q_{S}=\frac{L_{1}^{2}}{R^{2}}=\pi_{1}^{2} \rightarrow P_{1}^{2} \sim Q_{C}
$$

and

$$
x_{1}=R \frac{\cot \theta^{\prime}}{\sin \varphi^{\prime}} \rightarrow x, \quad x_{2}=R \cot \varphi^{\prime} \rightarrow y
$$

It is easy to see that for the case $a_{1}=a_{3}$ the corresponding spherical system of coordinates (1.64) contracts to Cartesian coordinates on $E_{2}$ for $R \rightarrow \infty$.
3. Elliptic coordinates on $S_{2}$ to elliptic coordinates on $E_{2}$. We take $Q$ in its general form, equivalent to

$$
\begin{equation*}
Q_{E}=L_{3}^{2}-\left(\frac{a_{2}-a_{1}}{a_{3}-a_{2}}\right) L_{1}^{2} \tag{3.1}
\end{equation*}
$$

We put

$$
\begin{equation*}
\frac{R^{2}}{a_{3}-a_{1}}=\frac{D^{2}}{a_{2}-a_{1}} \tag{3.2}
\end{equation*}
$$

and in the limit $R^{2} \sim a_{3} \rightarrow \infty$ obtain

$$
\begin{equation*}
Q_{E}=L_{3}^{2}-\left(\frac{a_{3}-a_{1}}{a_{3}-a_{2}}\right) \frac{D^{2}}{R^{2}} L_{1}^{2} \rightarrow L_{3}^{2}-D^{2} P_{2}^{2} \sim I_{E} \tag{3.3}
\end{equation*}
$$

For the coordinates we put

$$
\begin{equation*}
\rho_{1}=a_{1}+\left(a_{2}-a_{1}\right) \cos ^{2} \eta, \quad \rho_{2}=a_{1}+\left(a_{2}-a_{1}\right) \cosh ^{2} \xi \tag{3.4}
\end{equation*}
$$

and for $R^{2} \sim a_{3} \rightarrow \infty$, using Eq. (3.2), we obtain Eq. (3.41), i. e., elliptic coordinates on the plane $E_{2}$.
4. Elliptic coordinates on $S_{2}$ to Cartesian coordinates on $E_{2}$. We start from the coordinates (1.72) but change the ordering of the parameters $a_{i}$, which corresponds the interchange of coordinates $u_{3} \leftrightarrow u_{2}$, i.e., put

$$
\begin{equation*}
a_{1} \leq \rho_{1} \leq a_{3} \leq \rho_{2} \leq a_{2} \tag{3.5}
\end{equation*}
$$

and choose $a_{3}-a_{1}=a_{2}-a_{3} \equiv a$. Then we have

$$
\begin{equation*}
Q_{E}=a\left(L_{2}^{2}-L_{1}^{2}\right) \tag{3.6}
\end{equation*}
$$

Introducing the new coordinates by

$$
\begin{equation*}
\frac{a_{3}-\rho_{1}}{a}=\xi_{1}, \quad \frac{\rho_{2}-a_{3}}{a}=\xi_{2} \tag{3.7}
\end{equation*}
$$

we can rewrite the (1.72) in the form

$$
\begin{equation*}
u_{1}^{2}=\frac{R^{2}}{2}\left(1-\xi_{1}\right)\left(1+\xi_{2}\right), \quad u_{2}^{2}=\frac{R^{2}}{2}\left(1+\xi_{1}\right)\left(1-\xi_{2}\right), \quad u_{2}^{2}=R^{2} \xi_{1} \xi_{2} \tag{3.8}
\end{equation*}
$$

Using Eq. (2.1) we have for Beltrami coordinates

$$
\begin{equation*}
x_{1}^{2}=R^{2} \frac{\left(1-\xi_{1}\right)\left(1+\xi_{2}\right)}{2 \xi_{1} \xi_{2}}, \quad x_{2}^{2}=R^{2} \frac{\left(1+\xi_{1}\right)\left(1-\xi_{2}\right)}{2 \xi_{1} \xi_{2}} \tag{3.9}
\end{equation*}
$$

From equation (3.9) we obtain

$$
\begin{equation*}
\xi_{2,1}=\frac{R^{2}}{R^{2}+x_{1}^{2}+x_{2}^{2}}\left\{\left[1+\frac{x_{1}^{2}+x_{2}^{2}}{R^{2}}+\frac{\left(x_{1}^{2}-x_{2}^{2}\right)^{2}}{4 R^{4}}\right]^{1 / 2} \mp \frac{x_{1}^{2}-x_{2}^{2}}{2 R^{2}}\right\} . \tag{3.10}
\end{equation*}
$$

Taking now the limit $R \rightarrow \infty$ we have

$$
\begin{equation*}
\xi_{1} \rightarrow 1-\frac{x^{2}}{R^{2}}, \quad \xi_{2} \rightarrow 1-\frac{y^{2}}{R^{2}} \tag{3.11}
\end{equation*}
$$

and hence $x_{1}$ and $x_{2}$ of Eq. (3.9) go into Cartesian coordinates:

$$
\begin{equation*}
x_{1} \rightarrow x, \quad x_{2} \rightarrow y \tag{3.12}
\end{equation*}
$$

For the integral of motion in the limit $R^{2} \sim a \rightarrow \infty$ we have

$$
\begin{equation*}
\frac{1}{a R^{2}} Q_{E}=\left(\pi_{1}^{2}-\pi_{2}^{2}\right) \rightarrow P_{1}^{2}-P_{2}^{2}=Q_{C} \tag{3.13}
\end{equation*}
$$

5. Elliptic coordinates on $S_{2}$ to parabolic coordinates on $E_{2}$. We take the operator (1.72) with $a_{1} \leq \rho_{1} \leq a_{2} \leq \rho_{2} \leq a_{3}$ and choose the parameter
$a_{3}-a_{2}=a_{2}-a_{1} \equiv a$. We must first «undo» the diagonalization (1.26) by a rotation through $\pi / 4$. The operator (1.63) transforms into

$$
\begin{equation*}
\frac{1}{a R} Q_{E}=-\frac{1}{R}\left(L_{1} L_{3}+L_{3} L_{1}\right)=\left(L_{3} \pi_{2}+\pi_{2} L_{3}\right) \tag{3.14}
\end{equation*}
$$

with the correct limit (1.39) for $R \rightarrow \infty$. The coordinates (1.77) on $S_{2}$ are rotated into

$$
\begin{gather*}
u_{1}=\frac{R}{\sqrt{2}}(\operatorname{sn} \alpha \operatorname{dn} \beta+\operatorname{dn} \alpha \operatorname{sn} \beta), \quad u_{2}=R \operatorname{cn} \alpha \operatorname{cn} \beta \\
u_{3}=\frac{R}{\sqrt{2}}(\operatorname{dn} \alpha \operatorname{sn} \beta-\operatorname{sn} \alpha \operatorname{dn} \beta) \tag{3.15}
\end{gather*}
$$

with modulus $k=k^{\prime}=1 / \sqrt{2}$ for all Jacobi elliptic functions.
From Eq. (3.15) we obtain

$$
\begin{align*}
\operatorname{sn} \alpha & =\frac{1}{\sqrt{2}}\left[\left(1+\frac{u_{1}}{R}\right)^{1 / 2}\left(1-\frac{u_{3}}{R}\right)^{1 / 2}-\left(1-\frac{u_{1}}{R}\right)^{1 / 2}\left(1+\frac{u_{3}}{R}\right)^{1 / 2}\right] \\
\sqrt{2} \operatorname{dn} \beta & =\frac{1}{\sqrt{2}}\left[\left(1+\frac{u_{1}}{R}\right)^{1 / 2}\left(1-\frac{u_{3}}{R}\right)^{1 / 2}+\left(1-\frac{u_{1}}{R}\right)^{1 / 2}\left(1+\frac{u_{3}}{R}\right)^{1 / 2}\right] \tag{3.16}
\end{align*}
$$

Equations (3.16) suggest the limiting procedure. Indeed we put

$$
\begin{equation*}
\operatorname{sn} \alpha=-1+\frac{u^{2}}{2 R}, \quad \sqrt{2} \operatorname{dn} \beta=1+\frac{v^{2}}{2 R} \tag{3.17}
\end{equation*}
$$

In the limit $R \rightarrow \infty$ we obtain

$$
\begin{equation*}
x_{1} \rightarrow x=\frac{u^{2}-v^{2}}{2}, \quad x_{2} \rightarrow y=u v \tag{3.18}
\end{equation*}
$$

i. e., the parabolic coordinates (1.39).
3.2. Contractions of Coordinate Systems from $H_{2}$ to $E_{2}$

1. Pseudo-spherical coordinates on $H_{2}$ to polar coordinates on $E_{2}$. In the limit $R \rightarrow \infty, \tau \rightarrow 0$ putting $\tanh \tau \sim r / R$ we have:

$$
Q_{S}=L_{3}^{2} \rightarrow L_{3}^{2}
$$

and for Beltrami coordinates (2.7) we obtain:

$$
x_{1}=R \frac{u_{1}}{u_{0}} \rightarrow x=r \cos \varphi, \quad x_{2}=R \frac{u_{2}}{u_{0}} \rightarrow y=r \sin \varphi .
$$

2. Equidistant coordinates on $H_{2}$ to Cartesian on $E_{2}$. For Beltrami coordinates (2.7) we have:

$$
\begin{equation*}
x_{1}=R \tanh \tau_{2}, \quad x_{2}=R \tanh \tau_{1} / \cosh \tau_{2} \tag{3.19}
\end{equation*}
$$

Taking the limit $R \rightarrow \infty, \tau_{1}, \tau_{2} \rightarrow 0$ and putting $\sinh \tau_{1} \sim y / R, \sinh \tau_{2} \sim x / R$ in (3.19) we obtain $x_{1} \rightarrow x, x_{2} \rightarrow y$ and

$$
\frac{Q_{E Q}}{R^{2}}=\pi_{1}^{2} \rightarrow P_{1}^{2} \sim Q_{C}
$$

3. Horocyclic coordinates on $H_{2}$ to Cartesian on $E_{2}$. For variables $\tilde{x}, \tilde{y}$ we obtain:

$$
\tilde{x}=\frac{u_{2}}{u_{0}-u_{1}}, \quad \tilde{y}=\frac{R}{u_{0}-u_{1}}
$$

In the limit $R \rightarrow \infty$ we get: $\tilde{x} \rightarrow y / R, \quad \tilde{y} \rightarrow 1+x / R$ and Beltrami coordinates go into Cartesian ones

$$
x_{1}=R \frac{\tilde{x}^{2}+\tilde{y}^{2}-1}{\tilde{x}^{2}+\tilde{y}^{2}+1} \rightarrow x, \quad x_{2}=\frac{2 \tilde{x} R}{\tilde{x}^{2}+\tilde{y}^{2}+1} \rightarrow y
$$

For integral of motion we have:

$$
\frac{Q_{H O}}{R^{2}}=\pi_{2}^{2}+\frac{L_{3}^{2}}{R^{2}}-\frac{1}{R}\left\{\pi_{2}, L_{3}\right\} \rightarrow P_{2}^{2} \sim Q_{C}
$$

4. Elliptic coordinates on $\mathrm{H}_{2}$ to elliptic coordinates on $E_{2}$. We put

$$
\begin{equation*}
\frac{R^{2}}{a_{2}-a_{3}}=\frac{D^{2}}{a_{1}-a_{2}} \tag{3.20}
\end{equation*}
$$

and in the limit $R^{2} \sim\left(-a_{3}\right) \rightarrow \infty$ obtain:

$$
Q_{E}=L_{3}^{2}+\frac{D^{2}}{R^{2}} K_{2}^{2} \rightarrow L^{2}+D^{2} p_{1}^{2} \sim Q_{E}
$$

where $2 D$ is the focal distance. Writing the coordinates as

$$
\rho_{1}=a_{1}+\left(a_{1}-a_{2}\right) \sinh ^{2} \xi, \quad \rho_{2}=a_{2}+\left(a_{1}-a_{2}\right) \cos ^{2} \eta
$$

and using Eq. (3.20) in the limit $R^{2} \sim\left(-a_{3}\right) \rightarrow \infty$ we get the ordinary elliptic coordinates on $E_{2}$ plane [10, 15].
5. Elliptic coordinates on $H_{2}$ to Cartesian on $E_{2}$. We make a special choice of the parameters $a_{i}: a_{1}-a_{2}=a_{2}-a_{3}$ and determine new variables $\xi_{1,2}$ by the formula

$$
\begin{equation*}
\xi_{1,2}=\frac{\varrho_{1,2}-a_{2}}{a_{1}-a_{2}}=\frac{u_{0}^{2}+u_{2}^{2}}{2 R^{2}} \pm \sqrt{\left(\frac{u_{0}^{2}+u_{2}^{2}}{2 R^{2}}\right)^{2}-\frac{u_{1}^{2}}{R^{2}}} \tag{3.21}
\end{equation*}
$$

Considering the limit $R \rightarrow \infty$ we obtain: $\xi_{1} \sim 1+2 y^{2} / R^{2}, \xi_{2} \sim x^{2} / R^{2}$ and the Beltrami coordinate (2.7) takes the Cartesian form

$$
\begin{aligned}
& x_{1}=R \frac{u_{1}}{u_{0}}=R \sqrt{\frac{2 \xi_{1} \xi_{2}}{\left(\xi_{1}+1\right)\left(\xi_{2}+1\right)}} \rightarrow x \\
& x_{2}=R \frac{u_{2}}{u_{0}}=R \sqrt{\frac{\left(\xi_{1}-1\right)\left(1-\xi_{2}\right)}{\left(\xi_{1}+1\right)\left(\xi_{2}+1\right)}} \rightarrow y
\end{aligned}
$$

The operator $Q_{E}$ goes to Cartesian one

$$
\frac{Q_{E}}{R^{2}}=\frac{L_{3}^{2}}{R^{2}}+\pi_{1}^{2} \rightarrow P_{1}^{2} \sim Q_{C}
$$

6. Elliptic coordinates on $H_{2}$ to parabolic on $E_{2}$. We start from the rotated elliptic coordinates

$$
\begin{align*}
\left(\begin{array}{c}
u_{0}^{\prime} \\
u_{1}^{\prime} \\
u_{2}^{\prime}
\end{array}\right)=\left(\begin{array}{ccc}
\cosh f & \sinh f & 0 \\
\sinh f & \cosh f & 0 \\
0 & 0 & 1
\end{array}\right) & \left(\begin{array}{l}
u_{0} \\
u_{1} \\
u_{2}
\end{array}\right)= \\
& =\left(\begin{array}{c}
u_{0} \cosh f+u_{1} \sinh f \\
u_{0} \sinh f+u_{1} \cosh f \\
u_{2}
\end{array}\right) \tag{3.22}
\end{align*}
$$

where $\sinh ^{2} f=\left(a_{1}-a_{2}\right) /\left(a_{2}-a_{3}\right)$. We choose $a_{2}-a_{3}=a_{1}-a_{2} \equiv a$. Then for rotated elliptic coordiantes (3.22) we get

$$
\begin{gather*}
u_{0}=\frac{R}{\sqrt{2}}(\operatorname{sn} \alpha \operatorname{dn} \beta+i \sqrt{2} \operatorname{cn} \alpha \operatorname{cn} \beta)  \tag{3.23}\\
u_{1}=\frac{R}{\sqrt{2}}(i \operatorname{cn} \alpha \operatorname{cn} \beta+\sqrt{2} \operatorname{sn} \alpha \operatorname{dn} \beta), \quad u_{2}=i R \operatorname{dn} \alpha \operatorname{sn} \beta
\end{gather*}
$$

with modulus $k=k^{\prime}=1 / \sqrt{2}$ for all Jacobi elliptic function. The integral of motion transforms into

$$
\begin{equation*}
Q_{E^{\prime}}=3 L_{3}^{2}-\sqrt{2}\left(K_{1} L_{3}+L_{3} K_{1}\right) \tag{3.24}
\end{equation*}
$$

with the correct limit to (1.36). From Eq. (3.23) we obtain

$$
\begin{aligned}
& \operatorname{cn} \alpha=-\frac{i}{2} \sqrt{\left(1+\frac{u_{1}^{\prime}}{R \sqrt{2}}-\frac{u_{0}^{\prime}}{R}\right)^{2}+\frac{u_{2}^{\prime 2}}{2 R^{2}}}+\frac{i}{2} \sqrt{\left(1-\frac{u_{1}^{\prime}}{R \sqrt{2}}+\frac{u_{0}^{\prime}}{R}\right)^{2}+\frac{u_{2}^{\prime 2}}{2 R^{2}}} \\
& \operatorname{cn} \beta=\frac{1}{2} \sqrt{\left(1+\frac{u_{1}^{\prime}}{R \sqrt{2}}-\frac{u_{0}^{\prime}}{R}\right)^{2}+\frac{u_{2}^{\prime 2}}{2 R^{2}}}+\frac{1}{2} \sqrt{\left(1-\frac{u_{1}^{\prime}}{R \sqrt{2}}+\frac{u_{0}^{\prime}}{R}\right)^{2}+\frac{u_{2}^{\prime 2}}{2 R^{2}}}
\end{aligned}
$$

and therefore for large $R$ we have

$$
-i \operatorname{cn} \alpha \simeq 1-\frac{1}{2 \sqrt{2}} \frac{u^{2}}{R}, \quad \operatorname{cn} \beta \simeq 1+\frac{1}{2 \sqrt{2}} \frac{v^{2}}{R}
$$

In the limit $R \rightarrow \infty$ we obtain

$$
x_{1} \rightarrow x=\frac{u^{2}-v^{2}}{2}, \quad x_{2} \rightarrow y=u v
$$

i. e., the parabolic coordinates (1.39).
3.3. Contractions of Coordinate Systems from $H_{2}$ to $E_{1,1}$

1. Equidistant coordinates on $H_{2}$ to pseudo-spherical ones on $E_{1,1}$ plane. For Beltrami coordinates (2.10) we have:

$$
\begin{equation*}
y_{0}=R \operatorname{coth} \tau_{1} \cosh \tau_{2}, \quad y_{1}=R \operatorname{coth} \tau_{1} \sinh \tau_{2} \tag{3.25}
\end{equation*}
$$

Taking the limit $R \rightarrow \infty, \tau_{1} \rightarrow i \frac{\pi}{2}+\frac{r}{R}$ and putting

$$
\begin{equation*}
\operatorname{coth} \tau_{1}=\tanh \frac{r}{R} \sim \frac{r}{R} \tag{3.26}
\end{equation*}
$$

we obtain

$$
\begin{equation*}
y_{0} \rightarrow t=r \cosh \tau_{2}, \quad y_{1} \rightarrow x=r \sinh \tau_{2} \tag{3.27}
\end{equation*}
$$

where $0 \leq r<\infty,-\infty<\tau_{2}<\infty$. For the integral of motion we get

$$
\begin{equation*}
Q_{E Q}=K_{2}^{2} \rightarrow Q_{S}=L_{3}^{2} \tag{3.28}
\end{equation*}
$$

2. Pseudo-spherical coordinates on $H_{2}$ to Cartesian coordinates on $E_{1,1}$. For coordinates (2.10) we have

$$
\begin{equation*}
y_{0}=R \frac{\operatorname{coth} \tau}{\cos \varphi}, \quad y_{1}=R \cot \varphi \tag{3.29}
\end{equation*}
$$

Taking the limit $R \rightarrow \infty, \tau \rightarrow i \pi / 2, \varphi \rightarrow \frac{\pi}{2}$ and putting

$$
\begin{equation*}
\operatorname{coth} \tau \sim \frac{t}{R}, \quad \cot \varphi \sim \frac{x}{R} \tag{3.30}
\end{equation*}
$$

we see that Beltrami coordinates go into Cartesian ones

$$
\begin{equation*}
y_{0} \rightarrow t, \quad y_{1} \rightarrow x \tag{3.31}
\end{equation*}
$$

For the integral of motion we obtain

$$
\begin{equation*}
\frac{Q_{S}}{R^{2}}=\frac{L_{3}^{2}}{R^{2}} \rightarrow P_{2}^{2} \sim Q_{C} \tag{3.32}
\end{equation*}
$$

## 4. CONTRACTION OF BASIS FUNCTIONS ON $S_{2}$ AND $H_{2}$

### 4.1. Contraction of Spherical Basis and Interbasis Expansions

1. Spherical basis on $S_{2}$ to polar on $E_{2}$. We start from the standard spherical functions $Y_{l m}(\theta, \phi)$ as basis functions of irreducible representations of the group $O(3)$ (see, e. g., Ref. 37)

$$
\begin{align*}
& Y_{l m}(\theta, \phi)=(-1)^{(m+|m|) / 2}\left[\frac{2 l+1}{2} \frac{(l+|m|)!}{(l-|m|)!}\right]^{1 / 2} \frac{(\sin \theta)^{|m|}}{2^{|m|}|m|!} \times \\
& \quad{ }_{2} F_{1}\left(-l+|m|, l+|m|+1 ;|m|+1 ; \sin ^{2} \frac{\theta}{2}\right) \frac{\mathrm{e}^{i m \phi}}{\sqrt{2 \pi}} \tag{4.1}
\end{align*}
$$

In the contraction limit $R \rightarrow \infty$ we put

$$
\begin{equation*}
\tan \theta \sim \theta \sim \frac{r}{R}, \quad l \sim k R \tag{4.2}
\end{equation*}
$$

Using the asymptotic formulas

$$
\begin{align*}
\lim _{R \rightarrow \infty}{ }_{2} F_{1}\left(-k R, k R ;|m|+1 ; \frac{r^{2}}{4 R^{2}}\right) & ={ }_{0} F_{1}\left(|m|+1 ;-\frac{k^{2} r^{2}}{4}\right)  \tag{4.3}\\
\lim _{z \rightarrow \infty} \frac{\Gamma(z+\alpha)}{\Gamma(z+\beta)} & =z^{\alpha-\beta} \tag{4.4}
\end{align*}
$$

and formula

$$
\begin{equation*}
J_{\nu}(z)=\left(\frac{z}{2}\right)^{\nu} \frac{1}{\Gamma(\nu+1)}{ }_{0} F_{1}\left(\nu+1 ;-\frac{z^{2}}{4}\right) \tag{4.5}
\end{equation*}
$$

we obtain

$$
\begin{equation*}
\lim _{\substack{R \rightarrow \infty \\ \theta \rightarrow 0}} \frac{1}{\sqrt{R}} Y_{l m}(\theta, \phi)=(-1)^{(m+|m|) / 2} \sqrt{k} J_{|m|}(k r) \frac{\mathrm{e}^{i m \phi}}{\sqrt{2 \pi}} \tag{4.6}
\end{equation*}
$$

The result (4.6) is not new [37]. The point is that this asymptotic formula is obtained very naturally in the context of group contractions applied to the separation of variables.
2. Spherical basis on $S_{2}$ to Cartesian on $E_{2}$. We start from the coordinates $\left(\theta^{\prime}, \phi^{\prime}\right)$ in Eq. (1.64), but drop the primes, and write the corresponding spherical
functions as

$$
\begin{align*}
& Y_{l m}(\theta, \phi)=\frac{\sqrt{2 l+1}}{2 \pi} \mathrm{e}^{i m \phi}(\sin \theta)^{|m|} \times \\
\times & \left\{\begin{array}{l}
(-1)^{\frac{l+m}{2}}\left[\frac{\Gamma\left(\frac{l+m+1}{\Gamma}\right) \Gamma\left(\frac{l-m+1}{}\right)}{\Gamma\left(\left(\frac{l-m+2}{2}\right) \Gamma\left(\frac{l-m+2}{2}\right)\right.}\right]^{1 / 2}{ }_{2} F_{1}\left(-\frac{l-m}{2}, \frac{l+m+1}{2} ; \frac{1}{2} ; \cos ^{2} \theta\right) \\
(-1)^{\frac{l+m-1}{2}}\left[\frac{\Gamma\left(\frac{l+m+}{2}\right) \Gamma\left(\frac{l-m+2}{2}\right)}{\Gamma\left(\left(\frac{l+m+1}{2}\right) \Gamma\left(\frac{l-m+1}{2}\right)\right.}\right]^{1 / 2} 2 \cos \theta_{2} F_{1}\left(-\frac{l-m-1}{2}, \frac{l+m+2}{2} ; \frac{3}{2} ; \cos ^{2} \theta\right)
\end{array}\right. \tag{4.7}
\end{align*}
$$

for $l+m$ even and odd, respectively, we now put

$$
\begin{equation*}
l \sim k R, \quad m \sim k_{2} R, \quad \theta \sim \frac{\pi}{2}, \quad \phi \sim \frac{\pi}{2} \tag{4.8}
\end{equation*}
$$

and

$$
\begin{equation*}
\sin \theta \rightarrow 1, \quad \cos \theta \rightarrow \frac{x}{R}, \quad \cos \phi \rightarrow \frac{y}{R} \tag{4.9}
\end{equation*}
$$

The ${ }_{2} F_{1}$ hypergeometric functions simplify to ${ }_{0} F_{1}$ ones, the $\Gamma$ functions also simplify and the final result is that under the contraction we have

$$
\begin{align*}
\lim _{R \rightarrow \infty}(-1)^{-((l+m) / 2)} Y_{l m}(\theta, \phi) & =\sqrt{\frac{k}{k_{1}}} \frac{\mathrm{e}^{i k_{2} y}}{\pi}\left\{\begin{array}{c}
{ }_{0} F_{1}\left(\frac{1}{2} ; \frac{-k_{1}^{2} x^{2}}{4}\right) \\
-i\left(k_{1} x\right)_{0} F_{1}\left(\frac{3}{2} ; \frac{-k_{1}^{2} x^{2}}{4}\right)
\end{array}\right. \\
& =\sqrt{\frac{k}{k_{1}}} \frac{\mathrm{e}^{i k_{2} y}}{\sqrt{\pi}}\left\{\begin{array}{c}
\cos k_{1} x \\
-i \sin k_{1} x
\end{array}\right. \tag{4.10}
\end{align*}
$$

with $k_{1}{ }^{2}+k_{2}{ }^{2}=k^{2}$ and for $l+m$ even and odd, respectively. The parity properties of $Y_{l m}$ under the exchange $\theta \rightarrow \pi-\theta$ have led to the appearance of $\cos k_{1} x$ and $\sin k_{1} x$ in Eq. (4.10), instead of the usual Cartesian coordinate solution $\exp i\left(k_{1} x+k_{2} y\right)$.

Finally note that the factor $\sqrt{k / k_{1}}$ in formula (4.10) is connected with the contraction of Kronecker symbols to delta function

$$
\delta_{l l^{\prime}} \rightarrow \frac{1}{R} \delta\left(k-k^{\prime}\right)=\frac{1}{R} \frac{k}{k^{\prime}} \delta\left(k_{1}-k_{1}^{\prime}\right)
$$

3. Contraction in interbasis expansions. Let us now consider the contraction $R \rightarrow \infty$ for the interbasis expansion (1.69). The contraction of basis functions was presented in the formulas (4.6) and (4.10). In order to obtain the corresponding limit we need the asymptotic behavior of the «little» Wigner $d$ function for a
large $R$. It is easy to see that the expression of $d$ function for the angle $\pi / 2$ in terms of hypergeometric functions ${ }_{2} F_{1}$ (see for example [37]) is not applicable for the contraction limit when $\ell \rightarrow \infty$ and $m \rightarrow \infty$ simultaneously. To make this contraction we use an integral representation for the function $d_{m_{2}, m_{1}}^{l}(\beta)$ [37]

$$
\begin{align*}
d_{m_{2}, m_{1}}^{l}(\beta)= & \frac{i^{m_{2}-m_{1}}}{2 \pi}\left[\frac{\left(l+m_{2}\right)!\left(l-m_{2}\right)!}{\left(l+m_{1}\right)!\left(l-m_{1}\right)!}\right]^{1 / 2} \times \\
& \times \int_{0}^{2 \pi}\left(\mathrm{e}^{i \frac{\varphi}{2}} \cos \frac{\beta}{2}+i \mathrm{e}^{-i(\varphi / 2)} \sin \frac{\beta}{2}\right)^{\ell-m_{1}} \times \\
& \times\left(\mathrm{e}^{-i(\varphi / 2)} \cos \frac{\beta}{2}+i \mathrm{e}^{i(\varphi / 2)} \sin \frac{\beta}{2}\right)^{\ell+m_{1}} \mathrm{e}^{i m_{2} \varphi} d \varphi \tag{4.11}
\end{align*}
$$

which for the particular case of $\beta=\pi / 2$ can be presented in the following form

$$
\begin{align*}
d_{m_{2}, m_{1}}^{l}\left(\frac{\pi}{2}\right)=(-1)^{\left(l-m_{1}\right) / 2} & \frac{2^{l}}{\pi}\left[\frac{\left(l+m_{2}\right)!\left(l-m_{2}\right)!}{\left(l+m_{1}\right)!\left(l-m_{1}\right)!}\right]^{1 / 2} \times \\
& \times \int_{0}^{\pi}(\sin \alpha)^{l-m_{1}}(\cos \alpha)^{l+m_{1}} \mathrm{e}^{2 i m_{2} \alpha} d \alpha \tag{4.12}
\end{align*}
$$

Using now the formulas [44]

$$
\cos (2 n \alpha)=T_{n}(\cos 2 \alpha), \quad \sin (2 n \alpha)=\sin 2 \alpha U_{n-1}(\cos 2 \alpha)
$$

where $T_{l}(x)$ and $U_{l}(x)$ are Tchebyshev polynomials of the first and second kind. After integrating over $\alpha$, we obtain a representation of the Wigner $d$ function for angles $\pi / 2$ in terms of the hypergeometrical function ${ }_{3} F_{2}(1)$

$$
\begin{gather*}
d_{m_{2}, m_{1}}^{l}\left(\frac{\pi}{2}\right)=\frac{(-1)^{\frac{l-m_{1}}{2}}}{\sqrt{\pi} l!} \sqrt{\left(l+m_{2}\right)!\left(l-m_{2}\right)!} \times \\
\times\left\{\begin{array}{c}
\left\{\frac{\Gamma\left(\frac{l+m_{1}+1}{2}\right) \Gamma\left(\frac{l-m_{1}+1}{2}\right)}{\Gamma\left(\frac{l+m_{1}}{2}+1\right) \Gamma\left(\frac{l-m_{1}}{2}+1\right)}\right\}^{1 / 2}{ }_{3} F_{2}\left(\left.\begin{array}{c}
-m_{2}, m_{2}, \frac{l+m_{1}+1}{2} \\
\frac{1}{2}, l+1
\end{array}\right|_{1}\right),\left(l+m_{1}\right)-\text { even }, \\
\frac{2 i l}{(l+1)}\left\{\frac{\Gamma\left(\frac{l+m_{1}}{2}+1\right) \Gamma\left(\frac{l-m_{1}}{2}+1\right)}{\Gamma\left(\frac{l+m_{1}+1}{2}\right) \Gamma\left(\frac{l-m_{1}+1}{2}\right)}\right\}^{1 / 2}{ }_{3} F_{2}\left(\left.\begin{array}{c}
-m_{2}+1, m_{2}+1, \frac{l+m_{1}}{2}+1 \\
\frac{3}{2}, l+2
\end{array} \right\rvert\,\right),\left(l+m_{1}\right)-\text { odd. }
\end{array}\right. \tag{4.13}
\end{gather*}
$$

For large $R$ we put

$$
\begin{equation*}
l \sim k R, \quad m_{1} \sim k_{1} R, \quad \theta_{1} \sim \frac{r}{R}, \quad \theta_{1}^{\prime} \sim \frac{y}{R}, \quad \theta_{2}^{\prime} \sim \frac{x}{R} \tag{4.14}
\end{equation*}
$$

where $k^{2}=k_{1}^{2}+k_{2}^{2}$. Using the asimptotic formulas for ${ }_{3} F_{2}(1)$ function (4.3) and $\Gamma$ function (4.4), we get

$$
\begin{align*}
& \lim _{R \rightarrow \infty}(-1)^{-\frac{l-\left|m_{1}\right|}{2}} \sqrt{R} d_{m_{2}, m_{1}}^{l}\left(\frac{\pi}{2}\right)=\sqrt{\frac{2}{\pi k}} \times \\
& \times\left\{\begin{aligned}
\left(\frac{k^{2}}{k_{2}}\right)^{1 / 4}{ }_{2} F_{1}\left(-m_{2}, m_{2} ; \frac{1}{2} ; \frac{k+k_{1}}{2 k}\right) \\
-i m_{2}\left(\frac{k_{2}}{k^{2}}\right)^{1 / 4}{ }_{2} F_{1}\left(-m_{2}+1, m_{2}+1 ; \frac{3}{2} ; \frac{k+k_{1}}{2 k}\right)
\end{aligned}\right. \\
&=(-1)^{\left(3 m_{2}\right) / 2} \sqrt{\frac{2}{\pi k_{2}}}\left\{\begin{array}{l}
\cos m_{2} \varphi \\
i \sin m_{2} \varphi
\end{array}\right. \tag{4.15}
\end{align*}
$$

with $\cos \varphi=k_{1} / k$ and for $\left(\ell+m_{1}\right)$ even or odd, respectively.
Multiplying now the interbasis expansion (1.69) by the factor $(-1)^{-\frac{l-\left|m_{1}\right|}{2}}$, and taking the contraction limit $R \rightarrow \infty$ we obtain $\left(\theta \equiv \theta_{2}, m \equiv m_{2}\right)$

$$
\mathrm{e}^{i k_{1} x}\left\{\begin{array}{c}
\cos k_{2} y  \tag{4.16}\\
\sin k_{2} y
\end{array}\right\}=\sum_{m=-\infty}^{\infty}(i)^{|m|}\left\{\begin{array}{c}
\cos m \varphi \\
-\sin m \varphi
\end{array}\right\} J_{|m|}(k r) \mathrm{e}^{i m \theta}
$$

or in exponential form

$$
\begin{equation*}
\mathrm{e}^{i k r \cos (\theta-\varphi)}=\sum_{m=-\infty}^{\infty}(i)^{m} J_{m}(k r) \mathrm{e}^{i m(\theta-\varphi)} . \tag{4.17}
\end{equation*}
$$

The inverse expansion is

$$
\begin{equation*}
J_{m}(k r) \mathrm{e}^{i m \theta}=\frac{(-i)^{m}}{2 \pi} \int_{0}^{2 \pi} \mathrm{e}^{i m \varphi-i k r \cos (\theta-\varphi)} d \varphi \tag{4.18}
\end{equation*}
$$

For $\theta=0$ the two last formulas are equivalent to the well-known formulas in the theory of Bessel functions [44], namely expansions of plane waves in terms of cylindrical ones and vice versa.
4.2. Solutions of the Lamé Equation. Let us consider Eq. (1.60) on the sphere $S_{2}$ and separate variables in the elliptic coordinates (1.72). We obtain two ordinary differential equation of the form

$$
\begin{align*}
\frac{d^{2} \psi}{d \rho^{2}}+\frac{1}{2}\left\{\frac{1}{\rho-a_{1}}+\frac{1}{\rho-a_{2}}\right. & \left.+\frac{1}{\rho-a_{3}}\right\} \frac{d \psi}{d \rho}+ \\
& +\frac{1}{4}\left\{\frac{\lambda-l(l+1) \rho}{\left(\rho-a_{1}\right)\left(\rho-a_{2}\right)\left(\rho-a_{3}\right)}\right\} \psi=0 \tag{4.19}
\end{align*}
$$

or equivalently

$$
\begin{equation*}
4 \sqrt{P(\rho)} \frac{d}{d \rho} \sqrt{P(\rho)} \frac{d \psi}{d \rho}-\{l(l+1) \rho-\lambda\} \psi=0 \tag{4.20}
\end{equation*}
$$

where

$$
P(\rho)=\left(\rho-a_{1}\right)\left(\rho-a_{2}\right)\left(\rho-a_{3}\right)
$$

Equation (4.19) is the Lamé equation in algebraic form. It is a Fuchsian type equation with 4 regular singularities (at $a_{1}, a_{2}, a_{3}$, and $\infty$ ) [38-40,45-47].

Its general solution can be represented by a series expansion about any one of the singular points $a_{k}$ as

$$
\begin{equation*}
\psi(\rho)=\left(\rho-a_{1}\right)^{\alpha_{1} / 2}\left(\rho-a_{2}\right)^{\alpha_{2} / 2}\left(\rho-a_{3}\right)^{\alpha_{3} / 2} \sum_{t=0}^{\infty} b_{t}^{(k)}\left(\rho-a_{k}\right)^{t} \tag{4.21}
\end{equation*}
$$

where we have

$$
\alpha_{j}\left(\alpha_{j}-1\right)=0, \quad j=1,2,3
$$

and can choose $k$ equal to 1,2 , or 3 .
Substituting (4.21) into the Lamé equation (4.19) we obtain a three-term recursion relation for $b_{t}^{k}$

$$
\begin{align*}
\beta_{t}^{(k)} b_{t+1}^{(k)}+\left[\gamma_{t}^{(k)}+\lambda-l(l\right. & \left.+1) a_{k}\right] b_{t}^{(k)}+ \\
& +(2 t+\alpha-l-2)(2 t+\alpha+l-1) b_{t-1}^{(k)}=0 \tag{4.22}
\end{align*}
$$

with

$$
\begin{gather*}
\alpha=\alpha_{1}+\alpha_{2}+\alpha_{3}, \quad \alpha_{i k}=\alpha_{i}-\alpha_{k}, \quad b_{-1}=0 \\
\beta_{t}^{(k)}=4\left(a_{i}-a_{k}\right)\left(a_{j}-a_{k}\right)(t+1)\left(t+\alpha_{k}+1 / 2\right) \quad(i, j, k \text { cyclic })  \tag{4.23}\\
\gamma_{t}^{(k)}=-\left(a_{i}-a_{k}\right)\left(2 t+\alpha_{k}+\alpha_{j}\right)^{2}-\left(a_{j}-a_{k}\right)\left(2 t+\alpha_{k}+\alpha_{i}\right)^{2}
\end{gather*}
$$

The expansion (4.21) represents a Lamé function. Since we are interested in representations of $O(3)$, the sum in $\psi(\rho)$ must be a polynomial of order $N$, i. e., we must have

$$
\begin{equation*}
b_{N} \neq 0, \quad b_{N+1}=b_{N+2}=\ldots=0 \tag{4.24}
\end{equation*}
$$

for some $N$. The condition for this is that we have

$$
\begin{equation*}
l=2 N+\alpha \tag{4.25}
\end{equation*}
$$

and we obtain a secular equation for the eigenvalues $\lambda$, i. e., the separation constant in elliptic coordinates, by requiring that the determinant of the homogeneous linear system (3.14) for $\left\{b_{0}, b_{1}, \ldots, b_{N}\right\}$ should vanish. Since $N$ and $l$ must be integers, Eq. (4.25) implies that $\alpha$ and $l$ must have the same parity.

Numerous further properties of the Lamé polynomials, in the context of representations of the group $O(3)$, in the $O(3) \supset D_{2}$ basis, were established, e. g., in Refs. 38-40.

Here let us just represent the basis functions as

$$
\begin{equation*}
\Psi_{l \lambda}^{p q}\left(\rho_{1}, \rho_{2}\right)=A_{l \lambda}^{p q} \psi_{l \lambda}^{p q}\left(\rho_{1}\right) \psi_{l \lambda}^{p q}\left(\rho_{2}\right), \tag{4.26}
\end{equation*}
$$

where $A_{l \lambda}^{p q}$ is some normalization constant. The labels $p, q$ take values $\pm 1$ and identify representations of $D_{2}$. For each value of $l$ the values of $p, q$, and $\lambda$ label $(2 l+1)$ different states. Since the given representations $(p, q)$ of $D_{2}$ can figure more than once in the reduction of a representation of $O(3)$ corresponding to the given $l$, we are faced with a «missing label problem», resolved by the quantum number $\lambda$, i. e., the operator $Q$ of Eq. (1.63).

The expansions that we shall use for the Lamé polynomials in (4.26) are as in Eq. (4.21), but the summation over $t$ is from $t=0$ to $t=N$.
4.3. Elliptic Basis on $S_{2}$ to Cartesian Basis on $E_{2}$. We choose elliptic coordinates on $S_{2}$ as in Eq. (1.72), but with $a_{1}<a_{3}<a_{2}$, as in Eq. (3.5). We write the basis functions as in Eq. (4.26) with

$$
\begin{align*}
& \psi_{l \lambda}\left(\rho_{1}\right)=\left(\rho_{1}-a_{1}\right)^{\alpha_{1} / 2}\left(\rho_{1}-a_{2}\right)^{\alpha_{2} / 2}\left(\rho_{1}-a_{3}\right)^{\alpha_{3} / 2} \sum_{t=0}^{N} b_{t}^{(1)}\left(\rho_{1}-a_{1}\right)^{t} \\
& \psi_{l \lambda}\left(\rho_{2}\right)=\left(\rho_{2}-a_{1}\right)^{\alpha_{1} / 2}\left(\rho_{2}-a_{2}\right)^{\alpha_{2} / 2}\left(\rho_{2}-a_{3}\right)^{\alpha_{3} / 2} \sum_{t=0}^{N} b_{t}^{(2)}\left(\rho_{2}-a_{2}\right)^{t} \tag{4.27}
\end{align*}
$$

as in Eq. (4.21). The coefficients $b_{t}^{(j)}(j=1,2)$ satisfy the recursion relation (4.22) and we have $N=(l-\alpha) / 2$. We use the coordinates $\xi_{1}$ and $\xi_{2}$ introduced in Eq. (3.7) (for $a \equiv a_{3}-a_{1}=a_{2}-a_{3}$ ). Equation (4.27) reduces to

$$
\begin{gather*}
\psi_{l \lambda}\left(\xi_{1}\right)=(-1)^{\left(\alpha_{2}+\alpha_{3}\right) / 2} a^{\alpha / 2}\left(1-\xi_{1}\right)^{\alpha_{1} / 2}\left(1+\xi_{1}\right)^{\alpha_{2} / 2} \xi_{1}^{\alpha_{3} / 2} \sum_{t=0}^{N} C_{t}^{1}\left(1-\xi_{1}\right)^{t} \\
\psi_{l \lambda}\left(\xi_{2}\right)=(-1)^{\alpha_{2} / 2} a^{\alpha / 2}\left(1-\xi_{2}\right)^{\alpha_{2} / 2}\left(1+\xi_{2}\right)^{\alpha_{1} / 2} \xi_{2}^{\alpha_{3} / 2} \sum_{t=0}^{N} C_{t}^{2}\left(1-\xi_{2}\right)^{t} \tag{4.28}
\end{gather*}
$$

with $C_{t}^{(1)}=a^{t} b_{t}, C_{t}^{(2)}=(-a)^{t} b_{t}$. The recursion relations (4.22) now imply

$$
\begin{align*}
& 8(t+1)\left(t+\alpha_{1}+\frac{1}{2}\right) C_{t+1}^{(1)}+\left\{\mu^{(1)}-2\left(2 t+\alpha_{1}+\alpha_{3}\right)^{2}-\right. \\
& \left.-\left(2 t+\alpha_{1}+\alpha_{2}\right)^{2}\right\} C_{t}^{(1)}+(2 t+\alpha-l-2)(2 t+\alpha+l-1) C_{t-1}^{(1)}=0  \tag{4.29}\\
& -8(t+1)\left(t+\alpha_{2}+\frac{1}{2}\right) C_{t+1}^{(2)}+\left\{\mu^{(2)}++2\left(2 t+\alpha_{2}+\alpha_{3}\right)^{2}+\right. \\
& \left.\quad+\left(2 t+\alpha_{1}+\alpha_{2}\right)^{2}\right\} C_{t}^{(2)}-(2 t+\alpha-l-2)(2 t+\alpha+l-1) C_{t-1}^{(2)}=0
\end{align*}
$$

where

$$
\begin{equation*}
\mu^{(j)}=\frac{1}{a}\left[\lambda-a_{j} l(l+1)\right], \quad j=1,2 . \tag{4.30}
\end{equation*}
$$

The contraction limit is taken using Eq. (3.7) to relate $\xi_{1,2}$ to the Cartesian coordinates on $E_{2}$. Taking $l \sim k R$ we find

$$
\begin{equation*}
\mu^{(1)} \rightarrow 2 R^{2} k_{1}^{2}, \quad \mu^{(2)} \rightarrow-2 R^{2} k_{2}^{2}, \quad k=\sqrt{k_{1}^{2}+k_{2}^{2}} \tag{4.31}
\end{equation*}
$$

For $R \rightarrow \infty$ the recursion relations (4.29) simplify to two-term ones that can be solved to obtain

$$
\begin{equation*}
C_{t}^{(j)}=\frac{R^{2 t}}{\left(\alpha_{j}+1 / 2\right)_{t}}\left(\frac{-k_{j}^{2}}{4}\right)^{t} \frac{1}{t!} \tag{4.32}
\end{equation*}
$$

with

$$
\begin{aligned}
\left(\alpha_{j}+\frac{1}{2}\right)_{t}= & \left(\alpha_{j}+\frac{1}{2}\right)\left(\alpha_{j}+\frac{3}{2}\right) \ldots\left(\alpha_{j}-\frac{3}{2}+t\right) \\
& t \geq 1, \quad\left(\alpha_{j}+\frac{1}{2}\right)_{0}=1
\end{aligned}
$$

Substituting (4.3) into (4.28) we obtain

$$
\begin{gather*}
\psi_{l \lambda}\left(\xi_{1}\right)=(-1)^{\left(\alpha_{2}+\alpha_{3}\right) / 2} \frac{a^{\alpha / 2}}{R^{\alpha_{1}}} x^{\alpha_{1}}{ }_{0} F_{1}\left(\alpha_{1}+\frac{1}{2} ;-\frac{k_{1}^{2} x^{2}}{4}\right), \\
\psi_{l \lambda}\left(\xi_{2}\right)=(-1)^{\alpha_{2} / 2} \frac{a^{\alpha / 2}}{R^{\alpha_{2}}} y^{\alpha_{2}}{ }_{0} F_{1}\left(\alpha_{2}+\frac{1}{2} ;-\frac{k_{2}^{2} y^{2}}{4}\right) . \tag{4.33}
\end{gather*}
$$

Using now the formula (4.10) we find the contraction limit:

$$
\begin{align*}
A_{l \lambda}^{p q}(R) \psi_{l \lambda}\left(\xi_{1}, \xi_{2}\right) & \rightarrow A_{l \lambda}^{p q}(R) \psi_{k_{1}}(x) \psi_{k_{2}}(y)=A_{l \lambda}^{p q}(-1)^{\alpha_{3} / 2} a^{\alpha / 2} \times \\
& \times \begin{cases}\cos k_{1} x \cos k_{2} y, & \alpha_{1}=0, \alpha_{2}=0 \\
-\frac{1}{k_{2} R} \cos k_{1} x \sin k_{2} y, & \alpha_{1}=0, \alpha_{2}=1 \\
-\frac{1}{k_{1} R} \sin k_{1} x \cos k_{2} y, & \alpha_{1}=1, \alpha_{2}=0 \\
-\frac{1}{k_{1} k_{2} R^{2}} \sin k_{1} x \sin k_{2} y, & \alpha_{1}=1, \alpha_{2}=1\end{cases} \tag{4.34}
\end{align*}
$$

4.4. Elliptic Basis on $S_{2}$ to Elliptic Basis on $E_{2}$. Let us start from the elliptic coordinates (1.72) with $a_{1} \leq \rho_{1} \leq a_{2} \leq \rho_{2} \leq a_{3}$. We take the limit $R \rightarrow \infty, a_{3} \rightarrow \infty$ with $\sqrt{a_{3}} / R, a_{1}$ and $a_{2}$ finite. We introduce a constant $D$ as in Eq. (3.2). Elliptic coordinates on the plane $E_{2}$ are introduced via Eq. (3.4), so that the Cartesian coordinates $(x, y)$ are expressed in terms of the elliptic ones $(\xi, \eta)$ as in Eq. (1.41). Let us first take the limit in the separated equations (4.19). Going over to the variables $(\xi, \eta)$ from $\left(\rho_{1}, \rho_{2}\right)$ we obtain for $R \rightarrow \infty$ :

$$
\begin{align*}
& \frac{d^{2} \psi_{1}}{d \eta^{2}}+\left\{\mu-\frac{k^{2} D^{2}}{2}\left(\frac{a_{2}+a_{1}}{a_{2}-a_{1}}\right)-\frac{k^{2} D^{2}}{2} \cos 2 \eta\right\} \psi_{1}=0  \tag{4.35}\\
& \frac{d^{2} \psi_{2}}{d \xi^{2}}+\left\{\mu-\frac{k^{2} D^{2}}{2}\left(\frac{a_{2}+a_{1}}{a_{2}-a_{1}}\right)-\frac{k^{2} D^{2}}{2} \cosh 2 \xi\right\} \psi_{2}=0 \tag{4.36}
\end{align*}
$$

with

$$
\mu=\frac{\lambda}{a_{3}}, \quad l \sim k R
$$

In (4.35) we recognize the standard form of the Mathieu equation, whereas Eq. (4.36) is a modified Mathieu equation [41]. Thus, in the contraction limit, Lamé functions will go over into Mathieu ones. Moreover, periodic solutions of the Lamé equation go over into periodic solutions of Eq. (4.35).

The contraction limit can also be taken directly in the Lamé polynomials, using the expansion (4.21) (cut off at $t=N$ ). The result that we obtain is

$$
\begin{align*}
& \lim _{R \rightarrow \infty} \frac{\Psi_{l \lambda}\left(\rho_{1}\right)}{R^{\alpha_{3}}}=\left(a_{2}-a_{1}\right)^{\alpha / 2} \frac{(-1)^{\left(\alpha_{2}+\alpha_{3}\right) / 2}}{D^{\alpha_{3}}} \times \\
& \times(\cos \eta)^{\alpha_{1}}(\sin \eta)^{\alpha_{2}} \sum_{t=0}^{\infty} C_{t}(\cos \eta)^{2 t} \tag{4.37}
\end{align*}
$$

$$
\begin{align*}
\lim _{R \rightarrow \infty} \frac{\Psi_{l \lambda}\left(\rho_{2}\right)}{R^{\alpha_{3}}}=\left(a_{2}-a_{1}\right)^{\alpha / 2} & \frac{(-1)^{\alpha_{3} / 2}}{D^{\alpha_{3}}} \times \\
& \times(\cosh \xi)^{\alpha_{1}}(\sinh \xi)^{\alpha_{2}} \sum_{t=0}^{\infty} C_{t}(\cosh \xi)^{2 t} \tag{4.38}
\end{align*}
$$

where the expansion coefficients $C_{t}$ satisfy recursion relations obtained from Eq. (4.22), namely

$$
\begin{equation*}
4(t+1)\left(t+1 / 2+\alpha_{1}\right) C_{t}+\left\{\mu-\left(2 t+\alpha_{1}+\alpha_{2}\right)^{2}\right\} C_{t}-k^{2} D^{2} C_{t}=0 \tag{4.39}
\end{equation*}
$$

4.5. Elliptic Basis on $S_{2}$ to Parabolic Basis on $E_{2}$. Let us consider the contraction limit for the Lamé equations (4.19). To do this we use equations (1.76) with $a_{3}-a_{2}=a_{2}-a_{1}=a$, i. e., $k=k^{\prime}=1 / \sqrt{2}$, together with Eq. (3.17), to obtain

$$
\begin{equation*}
\rho_{1} \sim a_{1}+a\left(-1+\frac{u^{2}}{2 R}\right), \quad \rho_{2} \sim a_{1}+a\left(1+\frac{v^{2}}{R}\right) \tag{4.40}
\end{equation*}
$$

The equation (4.19) for $\rho=\rho_{1}$ and $\rho=\rho_{2}$ in the limit $R \rightarrow \infty$, with $l^{2} \sim k^{2} R^{2}$ and $\lambda-a_{2} l(l+1)=\mu R a$, yields the two equations

$$
\begin{equation*}
\frac{d^{2} \psi_{1}}{d u^{2}}+\left(k^{2} u^{2}+\mu\right) \psi_{1}=0, \quad \frac{d^{2} \psi_{2}}{d v^{2}}+\left(k^{2} v^{2}-\mu\right) \psi_{2}=0 \tag{4.41}
\end{equation*}
$$

respectively.
Thus the Lamé equations in the contraction limit go over into the equations (4.41) for parabolic cylinder functions [44]. The same is of course true for solutions. The expansion (4.21) is not suitable for the contraction limit. In view of Eq. (3.17) we need expansions in terms of the variables $(1+\operatorname{sn} \alpha)$ and $(1-\sqrt{2} \operatorname{dn} \beta)$. This is not hard to do, following for instance methods used in Ref. 48 to relate the wave functions of a two-dimensional hydrogen atom, calculated in different coordinate systems. The formulas are cumbersome, so we shall not present them here.

### 4.6. Contractions of Basis Functions from $H_{2}$ to $E_{2}$ and $E_{1,1}$

1. Pseudo-spherical basis on $H_{2}$ to polar basis on $E_{2}$. The pseudo-spherical eigenfunctions $\Psi_{\rho m}(\tau, \varphi)$ normalized to the Dirac delta-function, have the form:

$$
\begin{align*}
& \Psi_{\rho m}(\tau, \varphi)=\sqrt{\frac{\rho \sinh \pi \rho}{2 \pi^{2} R}}\left|\Gamma\left(\frac{1}{2}+i \rho+|m|\right)\right| \times \\
& \times P_{i \rho-1 / 2}^{|m|}(\cosh \tau) \exp (i m \varphi) \tag{4.42}
\end{align*}
$$

where $m=0, \pm 1, \pm 2, \ldots$ In the contraction limit $R \rightarrow \infty$ we put: $\tanh \tau \sim$ $\tau \sim r / R, \rho \sim k R$. Rewriting the Legendre function in terms of hypergeometric function as [49]

$$
\begin{aligned}
P_{i \rho-1 / 2}^{|m|}(\cosh \tau)= & \frac{\Gamma(1 / 2+i \rho+|m|)}{\Gamma(1 / 2+i \rho-|m|)} \frac{1}{|m|!2^{|m|}}{ }_{2} F_{1} \times \\
& \quad \times\left(\frac{1}{2}+|m|+i \rho, \frac{1}{2}+|m|-i \rho ; 1+|m| ;-\sinh ^{2} \frac{\tau}{2}\right) .
\end{aligned}
$$

Then using the asymptotic formula for hypergeometrical function ${ }_{2} F_{1}$ and $\Gamma$ function

$$
\begin{equation*}
\lim _{|y| \rightarrow \infty}|\Gamma(x+i y)| \exp \left(\frac{\pi}{2}|y|\right)|y|^{1 / 2-x}=\sqrt{2 \pi} \tag{4.43}
\end{equation*}
$$

we obtain in the contraction limit $R \rightarrow \infty$ :

$$
\lim _{R \rightarrow \infty} \Psi_{\rho m}(\tau, \varphi)=\sqrt{k} J_{|m|}(k r) \frac{\mathrm{e}^{i m \varphi}}{\sqrt{2 \pi}}
$$

2. Pseudo-spherical basis on $H_{2}$ to Cartesian basis on $E_{1,1}$. Taking the Legendre function in Eq. (4.42) in terms of two hypergeometric functions [49]

$$
\begin{aligned}
& P_{i \rho-1 / 2}^{m}(\cosh \tau)=\frac{\sqrt{\pi} 2^{m}(\sinh \tau)^{-m}}{\Gamma\left(\frac{3}{4}-\frac{m+i \rho}{2}\right) \Gamma\left(\frac{3}{4}-\frac{m-i \rho}{2}\right)} \times \\
& \times\left\{\begin{array}{c}
2 \cosh \tau \frac{\Gamma\left(\frac{3}{4}-\frac{m+i \rho}{2}\right) \Gamma\left(\frac{3}{4}-\frac{m-i \rho}{2}\right)}{\Gamma\left(\frac{1}{4}-\frac{m+i \rho}{2}\right) \Gamma\left(\frac{1}{4}-\frac{m-i \rho}{2}\right)} \times \\
\times{ }_{2} F_{1}\left(\frac{3}{4}-\frac{m+i \rho}{2}, \frac{3}{4}-\frac{m-i \rho}{2} ; \frac{3}{2} ; \cosh ^{2} \tau\right)+ \\
\quad+{ }_{2} F_{1}\left(\frac{1}{4}-\frac{m+i \rho}{2}, \frac{1}{4}-\frac{m-i \rho}{2} ; \frac{1}{2} ; \cosh ^{2} \tau\right)
\end{array}\right.
\end{aligned}
$$

Putting for large $R$

$$
\rho \sim k R, m \sim k_{1} R, \quad \operatorname{coth} \tau \sim \frac{t}{R}, \quad \cot \varphi \sim \frac{x}{R}, \quad k^{2}+k_{1}^{2}=k_{0}^{2}
$$

Using two asymptotic formulas

$$
\begin{aligned}
& \lim _{R \rightarrow \infty}{ }_{2} F_{1}\left(\frac{1}{4}-\frac{m+i \rho}{2}, \frac{1}{4}-\frac{m-i \rho}{2} ; \frac{1}{2} ; \cosh ^{2} \tau\right)= \\
& \quad={ }_{0} F_{1}\left(\frac{1}{2} ;-\frac{k_{0}^{2} t^{2}}{4}\right)=\cos \left(k_{0} t\right) \\
& \begin{aligned}
\lim _{R \rightarrow \infty}{ }_{2} F_{1}\left(\frac{3}{4}-\frac{m+i \rho}{2}, \frac{3}{4}-\frac{m-i \rho}{2} ; \frac{3}{2} ; \cosh ^{2} \tau\right) & = \\
& ={ }_{0} F_{1}\left(\frac{3}{2} ;-\frac{k_{0}^{2} t^{2}}{4}\right)=\frac{\sin \left(k_{0} t\right)}{k_{0} t}
\end{aligned}
\end{aligned}
$$

and formula (4.4) we finally obtain

$$
\begin{equation*}
\lim _{R \rightarrow \infty} \sqrt{R}|\Gamma(i \rho)| \Psi_{\rho m}(\tau, \varphi)=\sqrt{\frac{2}{k_{0}}} \mathrm{e}^{i k_{0} t-i k_{1} x} \tag{4.44}
\end{equation*}
$$

### 4.7. Contractions for Equidistant Basis on $\mathrm{H}_{2}$

1. Equidistant basis on $\mathrm{H}_{2}$ to Cartesian basis on $E_{2}$. In the equidistant system the normalized eigenfunctions $\Psi_{\rho \lambda}\left(\tau_{1}, \tau_{2}\right)$ have the form:

$$
\Psi_{\rho \lambda}\left(\tau_{1}, \tau_{2}\right)=\sqrt{\frac{\rho \sinh \pi \rho}{\cosh ^{2} \pi \lambda+\sinh ^{2} \pi \rho}}\left(\cosh \tau_{1}\right)^{-1 / 2} P_{i \lambda-1 / 2}^{i \rho}\left(-\tanh \tau_{1}\right) \mathrm{e}^{i \lambda \tau_{2}}
$$

To perform the contraction we write the Legendre function in terms of hypergeometric function [49]

$$
\begin{aligned}
& P_{i \lambda-1 / 2}^{i \rho}\left(-\tanh \tau_{1}\right)=\frac{\sqrt{\pi} 2^{i \rho}\left(\cosh \tau_{1}\right)^{-i \rho}}{\Gamma\left(\frac{3}{4}-a\right) \Gamma\left(\frac{3}{4}-b\right)} \times \\
& \times\left\{{ }_{2} F_{1}\left(\frac{1}{4}+a, \frac{1}{4}+b ; \frac{1}{2} ; \tanh ^{2} \tau_{1}\right)+2 \tanh \tau_{1} \frac{\Gamma\left(\frac{3}{4}-a\right) \Gamma\left(\frac{3}{4}-b\right)}{\Gamma\left(\frac{1}{4}-a\right) \Gamma\left(\frac{1}{4}-b\right)} \times\right. \\
& \left.\quad \times{ }_{2} F_{1}\left(\frac{3}{4}-a, \frac{3}{4}-b ; \frac{3}{2} ; \tanh ^{2} \tau_{1}\right)\right\}
\end{aligned}
$$

where $a=i(\rho-\lambda) / 2 ; b=i(\rho+\lambda) / 2$. For large $R$ we put $\rho \sim k R, \lambda \sim k_{1} R$; $\tau_{2} \sim x / R, \tau_{1} \sim y / R$, where $x, y$ are the Cartesian coordinates. Then using the
asymptotic formulas:

$$
\begin{gathered}
\lim _{R \rightarrow \infty}{ }_{2} F_{1}\left(\frac{1}{4}+a, \frac{1}{4}+b ; \frac{1}{2} ; \tanh ^{2} \tau_{1}\right)={ }_{0} F_{1}\left(\frac{1}{2} ;-\frac{y^{2} k_{2}^{2}}{4}\right)=\cos k_{2} y \\
\lim _{R \rightarrow \infty}{ }_{2} F_{1}\left(\frac{3}{4}-a, \frac{3}{4}-b ; \frac{3}{2} ; \tanh ^{2} \tau_{1}\right)={ }_{0} F_{1}\left(\frac{3}{2} ;-\frac{y^{2} k_{2}^{2}}{4}\right)=\frac{1}{k_{2} y} \sin k_{2} y
\end{gathered}
$$

where $k_{1}^{2}+k_{2}^{2}=k^{2}$, we finally get

$$
\lim _{R \rightarrow \infty} \Psi_{\rho \lambda}\left(\tau_{1}, \tau_{2}\right)=\sqrt{\frac{k}{\pi k_{2}}} \exp \left(i k_{1} x+i k_{2} y\right)
$$

2. Contraction from equidistant basis on $H_{2}$ to polar on $E_{1,1}$. Writting the Legendre function in terms of hypergeometric functions [49]

$$
\begin{aligned}
& \quad P_{i \lambda-1 / 2}^{i \rho}\left(\tanh \tau_{1}\right)=\frac{1}{\sqrt{2 \pi}}\left(\sinh \tau_{1}\right)^{i \rho}\left\{2^{-i \lambda}\left(\operatorname{coth} \tau_{1}\right)^{i \lambda+1 / 2} \frac{\Gamma(-i \lambda)}{\Gamma(1 / 2-i(\rho+\lambda))} \times\right. \\
& \times{ }_{2} F_{1}\left(\frac{1}{4}-\frac{i(\rho-\lambda)}{2}, \frac{3}{4}-\frac{i(\rho-\lambda)}{2} ; 1+i \lambda ; \operatorname{coth}^{2} \tau_{1}\right)+2^{i \lambda}\left(\operatorname{coth} \tau_{1}\right)^{-i \lambda+1 / 2} \times \\
& \left.\times \frac{\Gamma(i \lambda)}{\Gamma(1 / 2-i(\rho-\lambda)}{ }_{2} F_{1}\left(\frac{1}{4}-\frac{i(\rho+\lambda)}{2}, \frac{3}{4}-\frac{i(\rho+\lambda)}{2} ; 1-i \lambda ; \operatorname{coth}^{2} \tau_{1}\right)\right\} .
\end{aligned}
$$

Putting for large $R: \rho \sim k R$ and $\cosh \tau_{1} \sim r / R$, and using the asymptotic formulas for hypergeometric functions

$$
\begin{aligned}
\lim _{R \rightarrow \infty}{ }_{2} F_{1}\left(\frac{1}{4}-\frac{i(\rho-\lambda)}{2}, \frac{3}{4}-\frac{i(\rho-\lambda)}{2} ; 1+\right. & \left.i \lambda ; \operatorname{coth}^{2} \tau_{1}\right)= \\
& =\Gamma(1+i \lambda)\left(\frac{k r}{2}\right)^{-i \lambda} J_{i \lambda}(k r) \\
\lim _{R \rightarrow \infty}{ }_{2} F_{1}\left(\frac{1}{4}-\frac{i(\rho+\lambda)}{2}, \frac{3}{4}-\frac{i(\rho+\lambda)}{2} ; 1-\right. & \left.i \lambda ; \operatorname{coth}^{2} \tau_{2}\right)= \\
= & \Gamma(1-i \lambda)\left(\frac{k r}{2}\right)^{+i \lambda} J_{-i \lambda}(k r)
\end{aligned}
$$

we obtain

$$
\lim _{R \rightarrow \infty} \frac{1}{\sqrt{R}} \Psi_{\rho \lambda}\left(\alpha, \tau_{2}\right)=\sqrt{\frac{k}{2}} H_{i \lambda}^{(1)}(k r) \mathrm{e}^{i \lambda\left(\tau_{2}+i(\pi / 2)\right)}
$$

where $H_{\nu}^{(1)}(z)$ is the first kind of Hankel function.

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