# Modeling $\nu N$ interactions with account of the QCD corrections up to $O(\alpha_s^3)$ (S.Alekhin, IHEP, Protvino)

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## Validation of the factorization at low Q



- For the NOMAD data  $< Q^2 > \sim 5 \text{ GeV}^2$ , relatively low value
- The low-Q charged-leptons data are used to tune the PDFs and the high-twist terms
- The high-order QCD corrections, up to  $O(\alpha_s^3)$ , are taken into account

# The HO QCD corrections in DIS

Splitting Functions (up to  $O(\alpha_s^3)$ ):

(Moch-Vermasseren-Vogt 04)

Massless quarks coefficient functions (up to  $O(\alpha_s^3)$ )

(Zijlstra-van Neerven 91-92)

(Kazakov-Kotikov 92)

(Vermasseren-Moch-Vogt 05)

Heavy quarks coefficient functions (up to  $O(\alpha_s^2)$ ):

(Laenen-Riemersam-Smith-van Neerven 92-93)

## **Non-QCD** corrections

• The DIS structure functions are calculated using OPE

$$F_{2,\mathrm{T}}(x,Q) = F_{2,\mathrm{T}}^{\mathrm{LT}}(x,Q) + \frac{H_{2,\mathrm{T}}^{(2)}(x)}{Q^2} + \left(\frac{H_{2,\mathrm{T}}^{(4)}(x)}{Q^4}\right)$$

The leading-twist terms (entirely dominant at  $Q^2 \gtrsim 10 \text{ GeV}^2$ ). The twist-4 terms (contributes at  $Q^2 \lesssim 10 \text{ GeV}^2$ ) and the twist-6 terms (might contribute at  $Q^2 \lesssim 3 \text{ GeV}^2$ ) – no QCD evolution.

- The target-mass correction by Georgi-Politzer
- The deuteron nuclear corrections by Kulagin-Petti

# High-twist terms in the fit with $Q^2 > 1 \text{ GeV}^2$



- The HT terms in  $F_2$  demonstrate good convergence:  $H_2^{(4)}$  is much smaller than  $H_2^{(2)}$  and comparable to 0 within the errors.
- For  $F_{\rm T}$  the picture is different: the magnitudes of the twist-4 and twist-6 terms are comparable and somehow compensate each other (*poor convergence of the OPE?*)

#### Impact of the twist-6 terms on pulls of the fit



The twist-6 terms in  $F_{\rm T}$ arise due to mismatch of the SLAC and BCDMS data at  $Q^2 = 5 \div 10 \text{ GeV}^2$ and different y. The corrections of  $Q^2 > 1 \text{ GeV}^3$ to the coefficient functions do not help to resolve this discrepancy; the EW corrections seems not to be responsible too. In the final version of the fit twist-6 terms are set to 0.

## The HT terms of the final fit



The HT terms in  $F_2$  and  $F_T$  averaged over proton and neutron are very similar within the errors; therefore the HT term in  $F_L$  is comparable to 0. The constraint  $H_2^{lN} =$  $H_T^{lN}$  was further imposed everywhere.

7



The excess in SLAC data on R at  $x \sim 0.2$  with respect to the QCD predictions was considered as evidence of the big HT contribution to R (and  $F_{\rm L}$ )

(Miramontes-... 89)

Meanwhile this excess is evidently connected with the SLAC/BCDMS discrepancy and can be hardly attributed to the HT contribution.

# Gluons in the low-Q DIS fit



The change in G(x) due to the low-Q data is  $2-3\sigma$  at  $x \sim$ 0.2; this is correlated with the change in the structure function R. Other PDFs are less affected by the low-Q data.

# Extrapolation of the fit to Q = 0



- $F_2 \sim Q^2, F_L \sim Q^4$  at  $Q \to 0$ from the vector current conservation
- cubic spline interpolation between Q = 1 GeV and Q = 0.

# The Drell-Yan data kinematics



The Drell-Yan data are supplementary to the DIS ones.

# Impact of the DY data on the sea distribution

(sa-Melnikov-Petriello 06)



- Experimental errors in the sea is < 20 % at  $x \leq 0.7$ .
- The errors in PDFs due to variation of the DY scales are comparable to the experimental ones (the corrections of O(α<sup>2</sup><sub>s</sub>) by Anastasiou-Dixon-Melnikov-Petriello are crucial at this point ).

# Determination of the strange sea from the dimuon neutrino data (NuTeV and CCFR)

$$\nu_{\mu} + N \longrightarrow \mu^{-} + c + X$$

$$\hookrightarrow \mu^{+} + X$$

$W^+s \longrightarrow c$	$O(lpha_{ m s}^0)$
$W^+s \longrightarrow cg$	$O(lpha_{ m s}^1)$
$W^+g \longrightarrow c\overline{s}$	$O(lpha_{ m s}^1)$

(Gottschalk 81)







(data-fit)/fit



(data-fit)/fit



(data-fit)/fit

#### Strange sea asymmetry



- The NuTeV and CCFR data prefer asymmetry of different sign; averaging of both gives zero
- The MRST fit gives positive value, close to the NuTeV result
- The value of asymmetry is not very sensitive to the QCD, EW, and nuclear correction

## Total strange sea



- The strange suppression factor value from the combined fit is  $0.54 \pm 0.02$  at  $Q^2 = 20 \text{ GeV}^2$
- The CCFR analysis of their own data gives this value about 0.4; due to enhanced *d*-quark distribution defined from the inclusive sample

#### Status of the inclusive NuTeV data



The NuTeV data at  $x \sim 0.01$ go above the charged-leptons fit. This discrepancy cannot be removed due to account of the  $O(\alpha_s^3)$  corrections to the C-odd coefficient functions by Moch-Rogal-Vogt; modification of the nuclear corrections did not help too. At large x the NuTeV data also go above the the chargedleptons fit.

The HT terms in  $\nu N$  structure functions from the global fit including the CHORUS data



- $H_2^{\nu N} = H_T^{\nu N}$ , motivated my the chargedleptons fit
- $H_2^{\nu N}$  is in remarkable agreement to  $H_2^{lN}$ rescaled with the quarks charge
- $\int H_3^{\nu N}(x) dx$  is -0.10 ± 0.03 GeV<sup>2</sup>, in nice agreement to the early calculations by Braun-Kolesnichenko.

Uncertainty in the extraction of the Weinberg angle sine  $(s_W)$  due to PDFs

$$R^{\nu} = \frac{\sigma_{\rm NC}^{\nu}}{\sigma_{\rm CC}^{\nu}} \approx \frac{1}{2} - s_W^2 + \frac{5}{9}(1+r)s_W^4$$

$$r = \frac{\sigma_{\rm CC}^{\bar{\nu}}}{\sigma_{\rm CC}^{\nu}} \approx \frac{(U+D)/3 + U + D + 2S}{U+D+2S + (\bar{U}+\bar{D})/3} = 0.4999(24)$$

The uncertainty in r would lead to the uncertainty in  $s_W^2$  at the level of 0.00005, much better than expected experimental accuracy of the NOMAD analysis.

# Summary

- The fit of the PDFs and HTs to the combined charged leptons DIS, fixed-target Drell-Yan, dimuon neutrino data by NuTeV and CCFR, and inclusive neutrino data by CHORUS demonstrates reasonable consistency of the data:  $\chi^2/NDP = 5177/4338 = 1.2; \ \alpha_s(M_Z) = 0.1138(7).$
- The charged leptons DIS data are well described down to  $Q^2 = 1 \ GeV^2$  with account of the QCD corrections up to  $O(\alpha_s^3)$ and down to  $Q^2 \approx 0.5 \ GeV^2$  using the spline interpolation combined with the current-conservation constraints.
- The HT terms extracted from the fit demonstrate remarkable universality:  $H_2^{lN} \approx H_T^{lN} \approx 5/18 H_{2,T}^{\nu N}$ .
- Accuracy of the PDFs obtained is quite sufficient for the precise extraction of the Weinberg angle from the NOMAD data.