

NEUTRINO INTERACTIONS WITH NUCLEAR MATTER AT LOW ENERGIES AND WITH NUCLEONS AT HIGH ENERGIES

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OUTLOOK

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VII. *Inclusive and semi-inclusive $\nu - A \rightarrow \mu^- + X$ processes*

VIII. *Summary (**low & high energies**)*

The main goal is to construct a new generator simulating neutrino interaction within a detector

*G.I.Lykasov, V.A.Bednyakov, Phys.Rev. C **014622** (2007); G.I.Lykasov, E.Olsson, C.J.Pethick, Phys.Rev.C **72** 025805 (2005); G.I.Lykasov, C.J.Pethick, A.Schwenk, Phys.Rev.C **78** 045803 (2008).*

The main inputs are the neutrino-nucleus cross sections in a wide energy region.

The rate of neutrino-nucleon scattering in a medium at low energies can be presented in the following form

$$W_{fi} = \frac{G_F^2 n}{4V} [C_V^2 (1 + \cos \theta) \mathcal{S}_V(\mathbf{q}, \omega) + C_A^2 (3 - \cos \theta) \mathcal{S}_A(\mathbf{q}, \omega)]$$

where θ is the scattering angle, V is the normalized volume, n is the nuclear density.

Fermi-liquid theory and neutrino scattering of nuclear matter

The FF $\mathcal{S}_{V,A}$ are related to the corresponding response function $\chi_{V,A}$

$$\mathcal{S}_{V,A}(\omega, \mathbf{q}) = \frac{2}{n} \frac{\text{Im}\chi_{V,A}(\omega, \mathbf{q})}{1 - \exp(-\omega/T)}.$$

The Dyson type perturbation equations over the spin-independent \mathcal{F} and spin dependent \mathcal{G} interactions of quasiparticles presented in the matrix form.

$$\begin{aligned}\chi_V &= \chi^0 - \chi_V \mathcal{F} \chi^0, \\ \chi_A &= \chi^0 - \chi_A \mathcal{G} \chi^0,\end{aligned}$$

Here χ^0 is the diagonal 2×2 matrix consisting of χ_p^0 and χ_n^0 which being the zero approximations of the proton and neutron response functions over the interaction. For isospin-symmetric nuclear matter \mathcal{F} and \mathcal{G} become also 2×2 matrices

$$\begin{aligned}\chi_V^p (1 + f_{nn} \chi_n^0) + \chi_V^n f_{pn} \chi_p^0 &= \chi_p^0 \\ \chi_V^p f_{pn} \chi_n^0 + \chi_V^n (1 + f_{pp} \chi_n^0) &= \chi_n^0\end{aligned}$$

and

$$\begin{aligned}\chi_A^p (1 + g_{nn} \chi_n^0) + \chi_A^n g_{pn} \chi_p^0 &= \chi_p^0 \\ \chi_A^p g_{pn} \chi_n^0 + \chi_A^n (1 + g_{pp} \chi_n^0) &= \chi_n^0,\end{aligned}$$

where f_{pp} , f_{nn} , f_{pn} and g_{pp} , g_{nn} , g_{pn} are the spin-independent and spin-dependent amplitudes of pp , nn and pn interactions, respectively.

Note, that the amplitude of interaction between two quasi-particles q and q' with three-momenta \mathbf{p}

and \mathbf{p}' neglecting the tensor forces has the following form

$$f_{qq'}(\mathbf{p}, \mathbf{p}') = f + f'(\boldsymbol{\tau} \cdot \boldsymbol{\tau}') + g(\boldsymbol{\sigma} \cdot \boldsymbol{\sigma}') + g'(\boldsymbol{\sigma} \cdot \boldsymbol{\sigma}')(\boldsymbol{\tau} \cdot \boldsymbol{\tau}')$$

where q and q' can denote p, n , and f, f', g, g' are the Landau parameters, $\boldsymbol{\sigma}$ and $\boldsymbol{\tau}$ are the spin and isospin Pauli matrices, respectively.

$$\begin{aligned} f_{pp} &= f_{nn} = f + f', \\ g_{pp} &= g_{nn} = g + g', \\ f_{pn} &= f_{np} = f - f', \\ g_{pn} &= g_{np} = g - g'. \end{aligned}$$

L.D.Landau, Sov.Phys.JETP, 5, 101(1957)

A.B.Migdal, "Theory of Finite Fermi Systems and Application to Atomic Nuclei", Interscience, New York, 1962

G.Baym, C.J.Pethick, "Landau Fermi-Liquid Theory: Concepts and Applications", New York, 1991

N.Iwamoto, C.J.Pethick, Phys.Rev. D25, 313 (1982)

G.L., E.Olsson, C.J.Pethick, Phys.Rev. C72, 02805 (2005)

S.Reddy, M.Prakash & J.M.Lattimer, Phys.Rev. C59, 2888 (1999)

Application of the FLT to the analysis of $\nu - A$ interactions

$$1/l = V \int \frac{d^3q}{(2\pi)^3} W_{fi}.$$

With this quantity one can estimate the cross section of the elastic neutrino interaction with a heavy nucleus σ_{el}

$$\sigma_{el} = \frac{V_A}{l} = V_A \int \frac{d^3q}{(2\pi)^3} \tilde{W}_{fi}$$

where $\tilde{W}_{fi} = V_A \cdot W_{fi}$ and $V_A = A \cdot v_N$. Here A is the number of nucleons in a nucleus and $v_N = 4\pi/3r_N^3$ is the nucleon volume, r_N is the nucleon radius about 0.8 fm. To estimate the number of neutrino interactions \mathcal{R} per 1 second within a target T we use the simple formula

$$\mathcal{R} = P_{targ} N_A \sigma_{\nu A} f_{\nu}$$

Here f_{ν} denotes the initial neutrino flux.

Comparison of the obtained results with other calculations

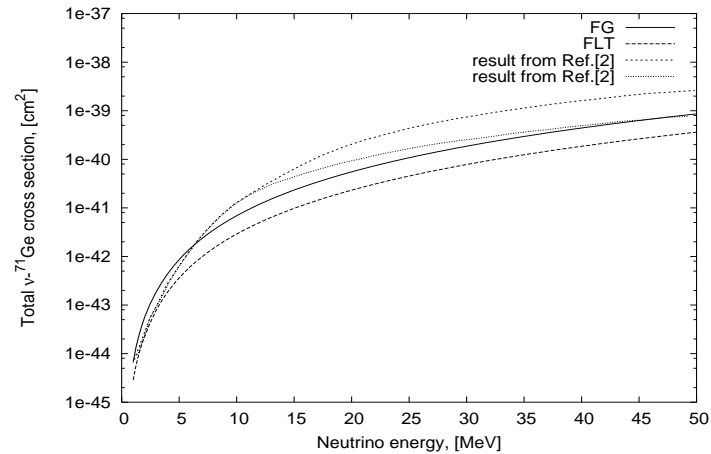
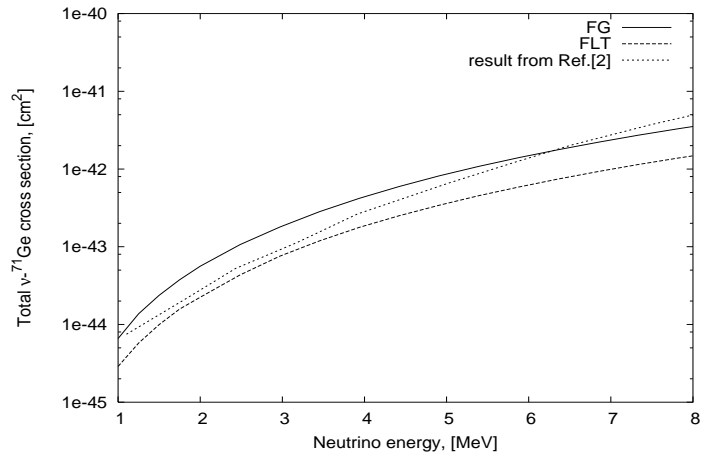


Figure 1: The total ν - ^{71}Ge cross section as a function of the neutrino energy E_ν .

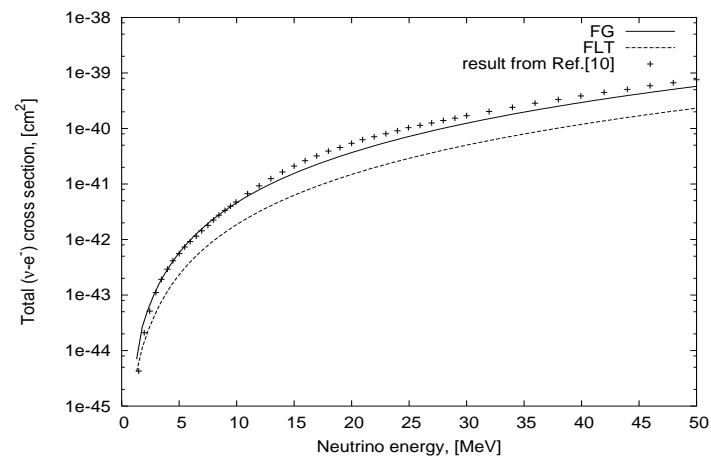
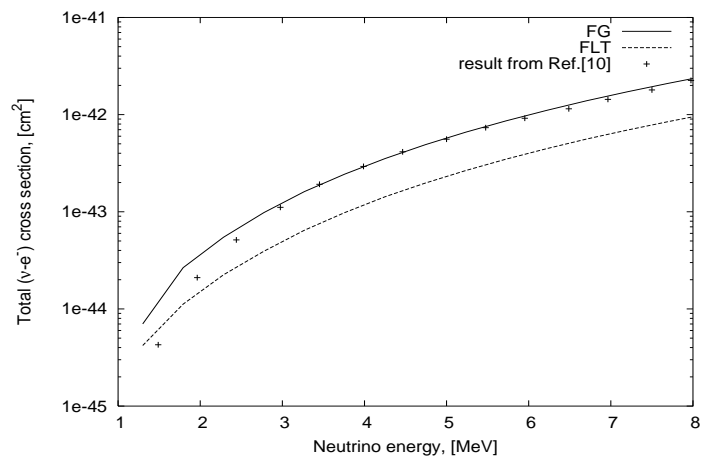


Figure 2: The total absorption ν - ^{40}Ar cross section as a function of the neutrino energy E_ν .

Background from solar neutrinos

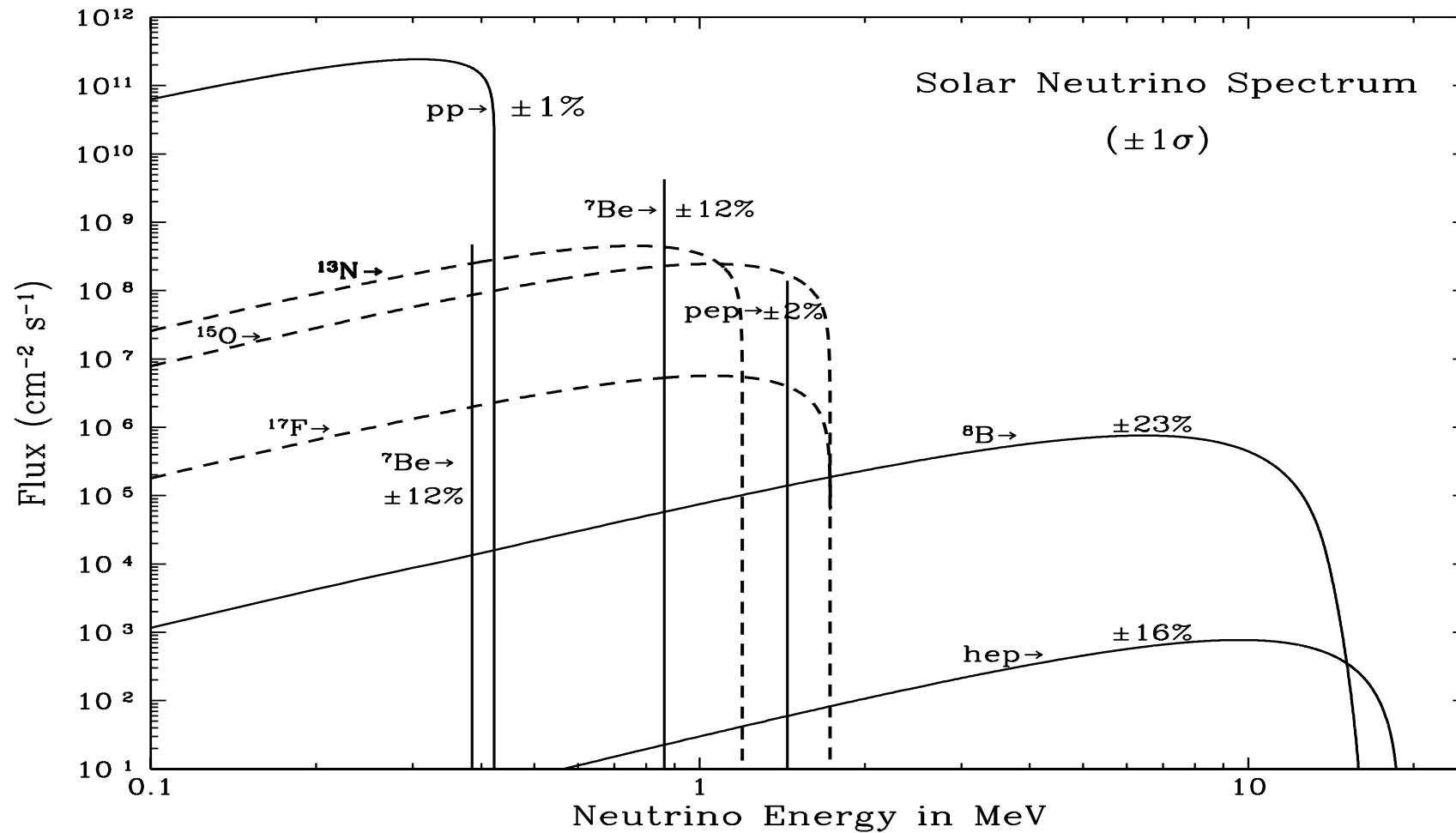


Figure 3: The flux continuum [$\text{cm}^{-2} \text{sec}^{-1} \text{MeV}^{-1}$] as a function of the neutrino energy E_ν .

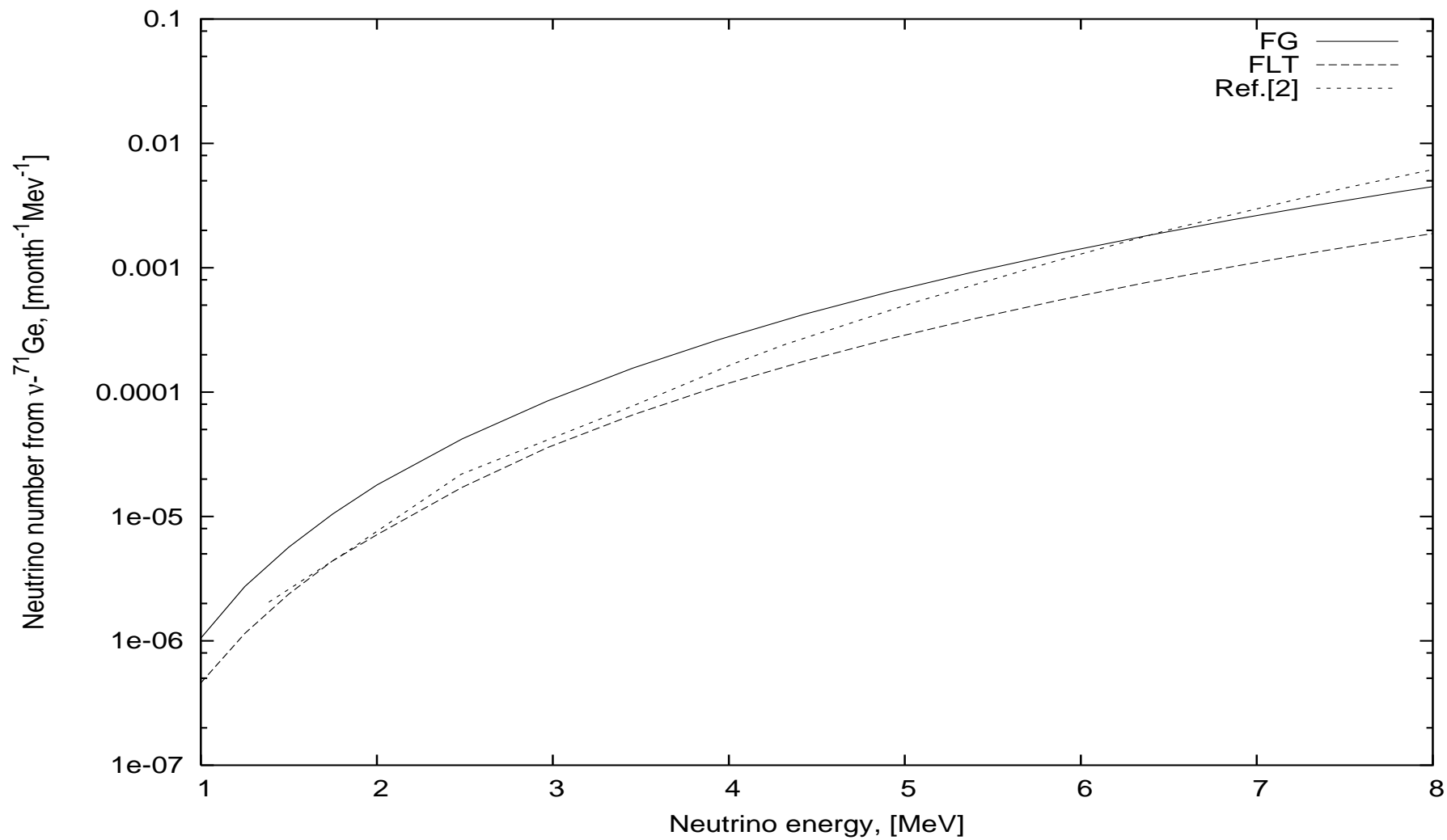


Figure 4: The total neutrino number per a month and MeV produced from ${}^8\text{B}-\nu$ flux interacting with 1.0 kg ${}^{71}\text{Ge}$ target as a function of the neutrino energy E_ν .

High energies

G.I.Lykasov, U.Sukhatme, V.V.Uzhinsky, *Phys.Lett. B* **553** 217 (2003); O.Benhar, S.Fantoni, G.I.Lykasov, U.Sukhatme, V.V.Uzhinsky, *Eur.Phys.J., A* **19** 147 (2004); O.Benhar, S.Fantoni, G.I.Lykasov, U.Sukhatme, *Phys.Lett. B* **527** 73 (2002); O.Benhar, S.Fantoni, G.I.Lykasov, *Eur.Phys.J. A* **7** 415 (2000).

Planar and cylinder graphs

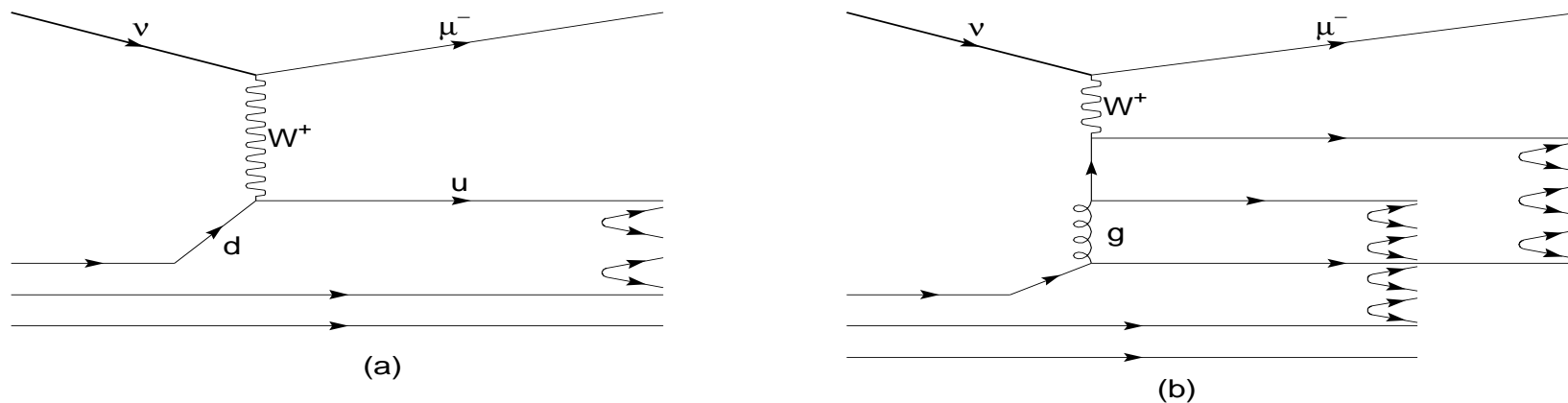


Figure 5: The planar graph (left panel) and the cylinder graph (right panel). *G.Veneziano, Phys.Lett., B52,220 (1974)*

Semi-inclusive process $\nu(\bar{\nu})p \rightarrow \mu^-(\mu^+)hX$

$$\rho_{nu(\bar{\nu})p \rightarrow \mu^-(\mu^+)hX} = \Phi(Q^2) \{ F_P(x, Q^2; z, p_t) + F_C(x, Q^2; z, p_t) \}$$

with

$$\Phi(Q^2) = mE \frac{G^2}{\pi} \frac{m_W^2}{Q^2 + m_W^2},$$

where G is the Fermi weak coupling constant, E is the energy of incoming neutrino, m and m_W are the nucleon and the W -boson masses respectively. $x = Q^2/2(p_\nu \cdot k)$ is the Bjorken variable, p_ν and k are the four-momenta of the initial neutrino and nucleon, $z = (E_h + p_{hz})/(E + p_z)$ is the light cone variable.

The variable z can be treated also as the Feynman variable $x_F = \frac{2p_L^*}{W_X}$ defined as the longitudinal momentum fraction in the hadronic center mass system (h.c.m.s.), p_L^* is the longitudinal hadron momentum in the h.c.m.s. (*G.L., U.Sukhatme, V.V.Uzhinsky, Phys.Lett.* **B553**,217 (2003))

Mean multiplicity of charged hadron in the current fragmentation region

The multiplicity $\langle n_{ch} \rangle$ measured by NOMAD is close to $\langle n_{ch} \rangle / 2$ results from e^+e^- experiment at $E = \sqrt{s}$ and $\langle n_{ch} \rangle$ from ep and $\bar{\nu}p$ at $E = Q$. **QCD fit for $\langle n_{ch} \rangle$**

$$\langle n_{ch}^{QCD} \rangle = a + b \exp(c \sqrt{\ln(Q^2/Q_0^2)}) ,$$

where $a=2.257$, $b=0.094$, $c=1.775$, $Q_0=1\text{GeV.c.}$ (*W.Furmanski, R.Petronzio, S.Pokorski, Nucl.Phys.***B155**,253 (1979); *A.Bassetto, M.Ciafaloni, G.Marchesini, Phys.Lett.***B83**, 207(1979); *K.Konishi, Rutherford Report RL 79-035 (1979); A.H.Mueller, Phys.Lett.***B213**, 85(1983))

Multiple hadron production in $\nu - p \rightarrow \mu^- + X$ process

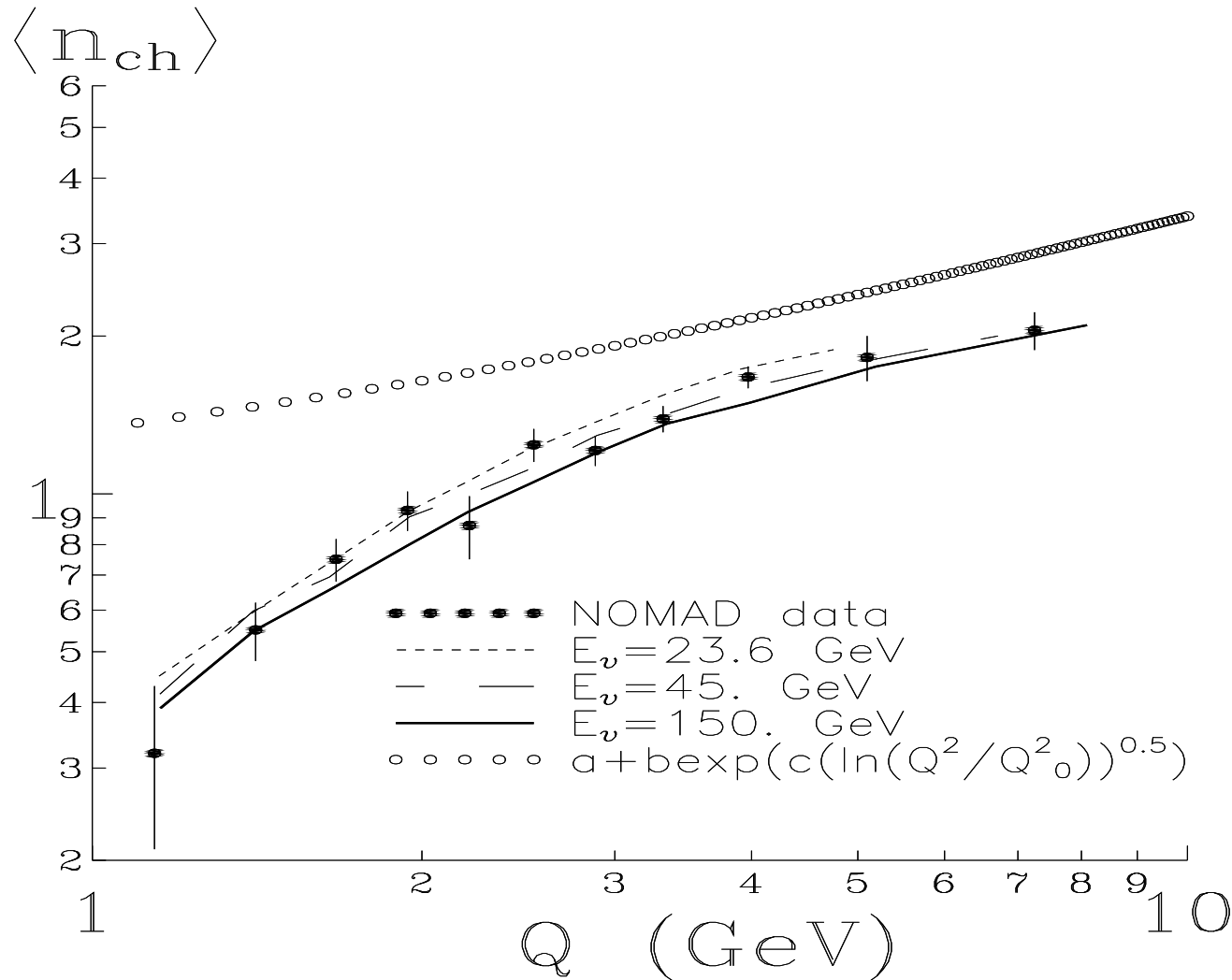


Figure 6: Mean multiplicity of charged hadron in the current fragmentation region as a function of the momentum transfer Q . The open circles correspond to the QCD fit; the solid, long dash and short dash lines correspond to our calculations at $E_\nu = 150$.GeV, 45.GeV and $E_\nu = 23.6$.GeV respectively. The experimental points are the NOMAD data.

Q^2 -inclusive spectrum

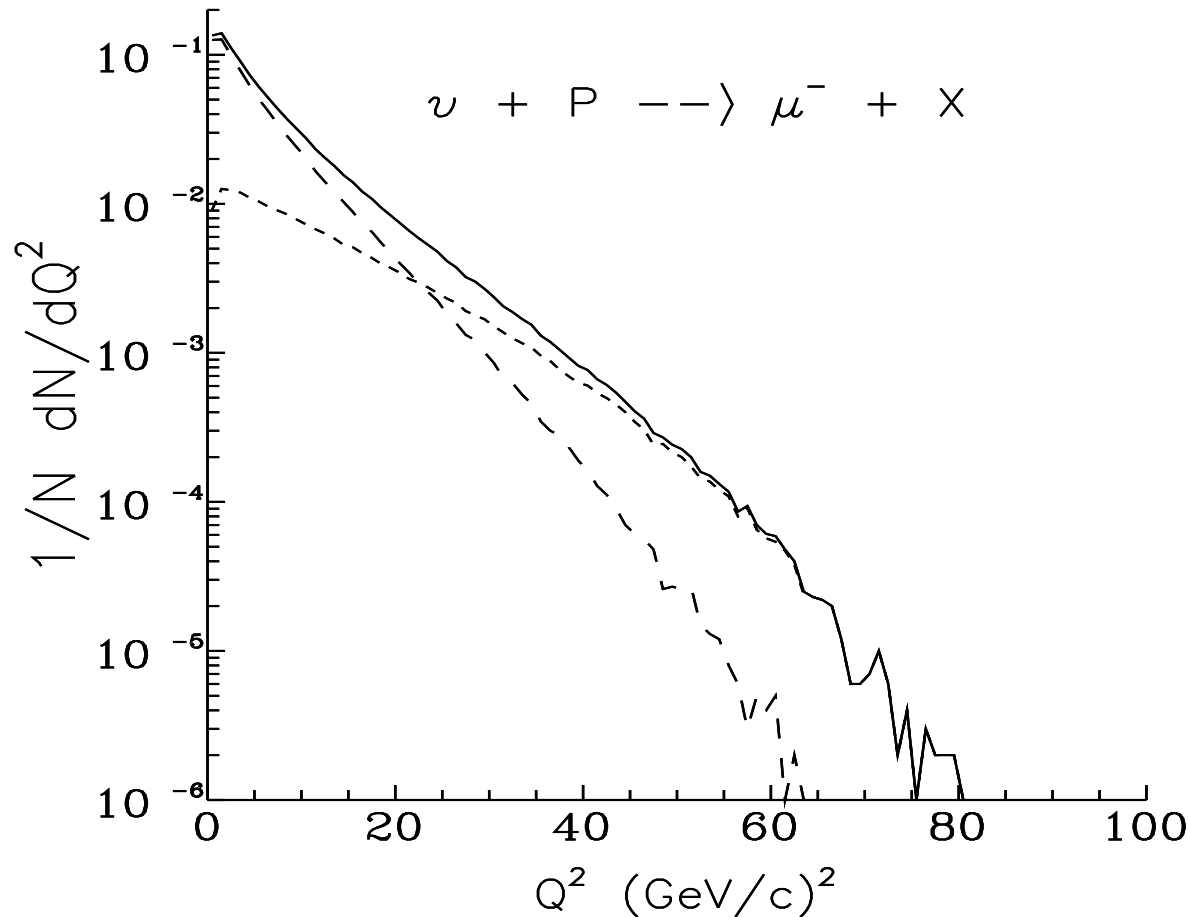


Figure 7: The Q^2 -distribution $\frac{1}{N} \frac{dN}{dQ^2}$ of muons produced in $\nu p \rightarrow \mu^- X$ reaction. The long dash and short dash lines correspond to the contributions of the cylinder and planar graphs respectively. The solid line is the sum of these contributions.

x_F -distributions of strange hadron

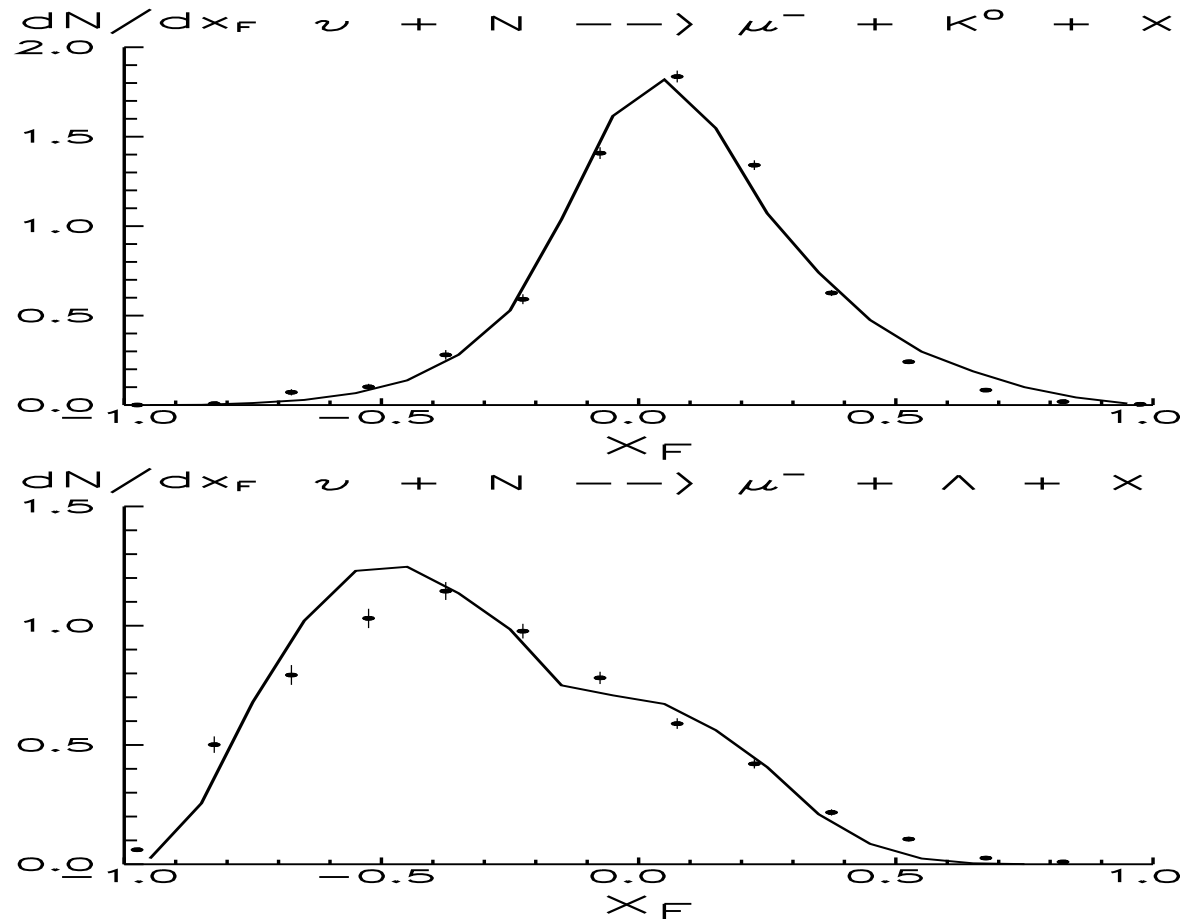


Figure 8: The x_f -distribution of strange hadron DN/dx_F . produced in $\nu p \rightarrow \mu^- X$ reaction.

Multiplicity of strange hadron

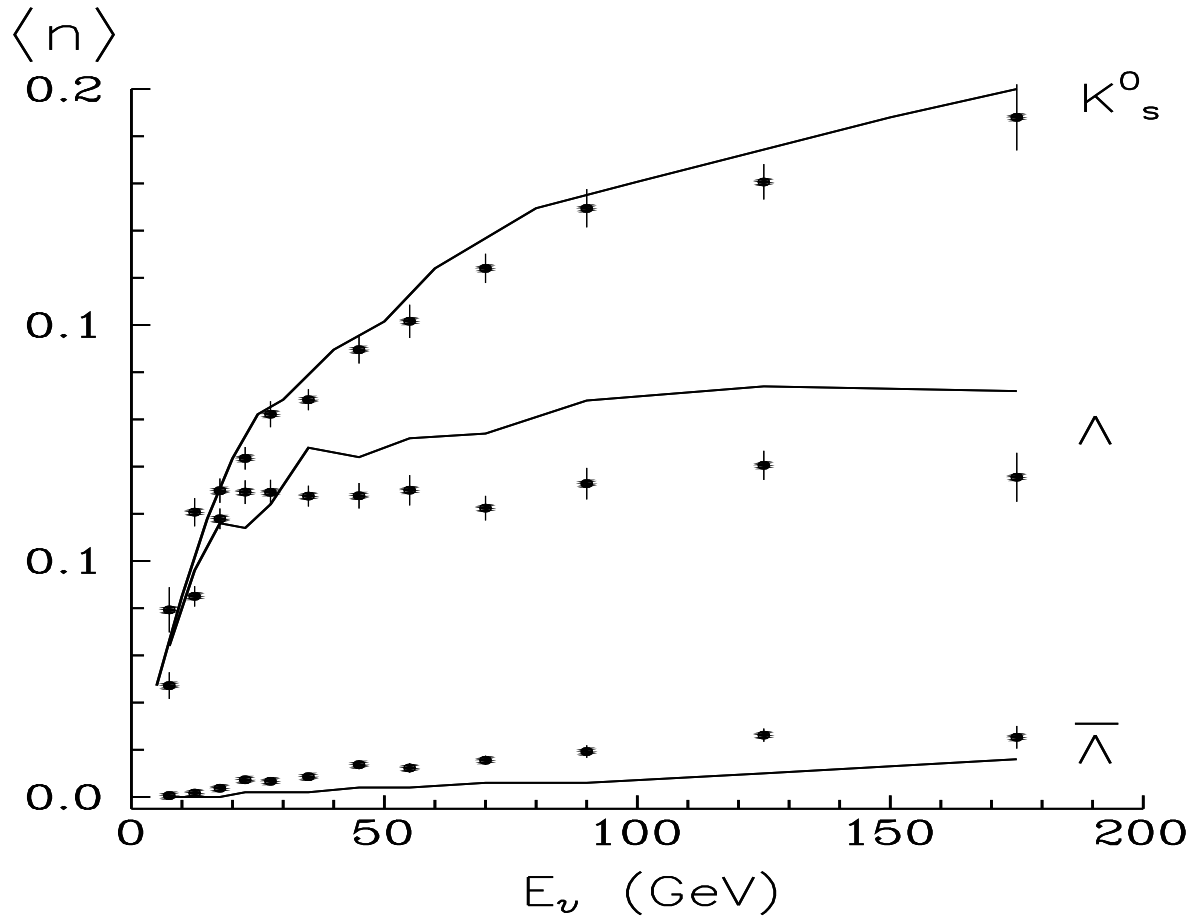


Figure 9: The multiplicity of strange hadron as a function of neutrino energy E_ν .

0 50 100 150 2

Multiplicity of backward going charged pions

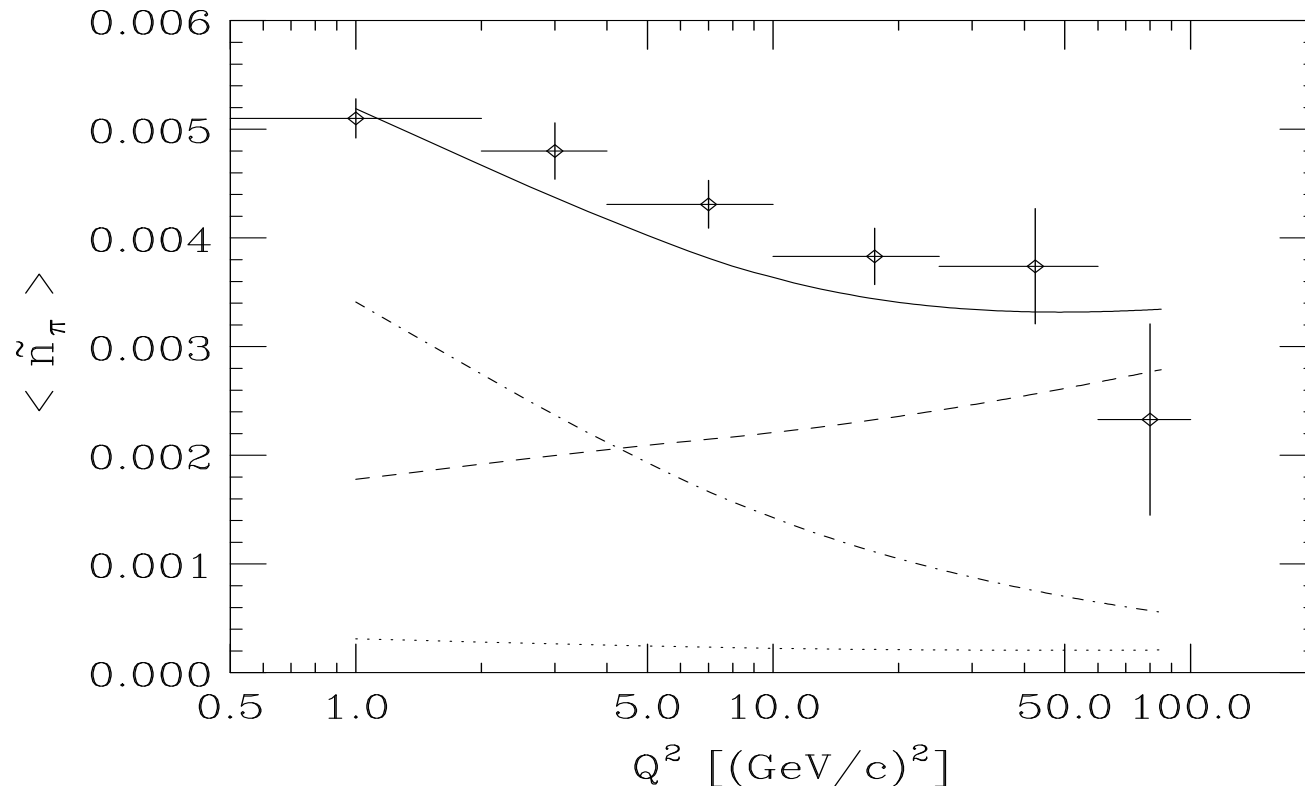


Figure 10: Mean multiplicity of charged pions produced in backward semi-sphere in $\nu^{12}C \rightarrow \mu^- \pi X$ process

(*P.Astier, et al., (NOMAD Coll.) Nucl.Phys..B609,255 (2001)*)

Our calculations

(*O.Benhar,S.Fantoni, G.L., U.Sukhatme, Phys.Lett.B527,73 (2002)*)

Backward going pions

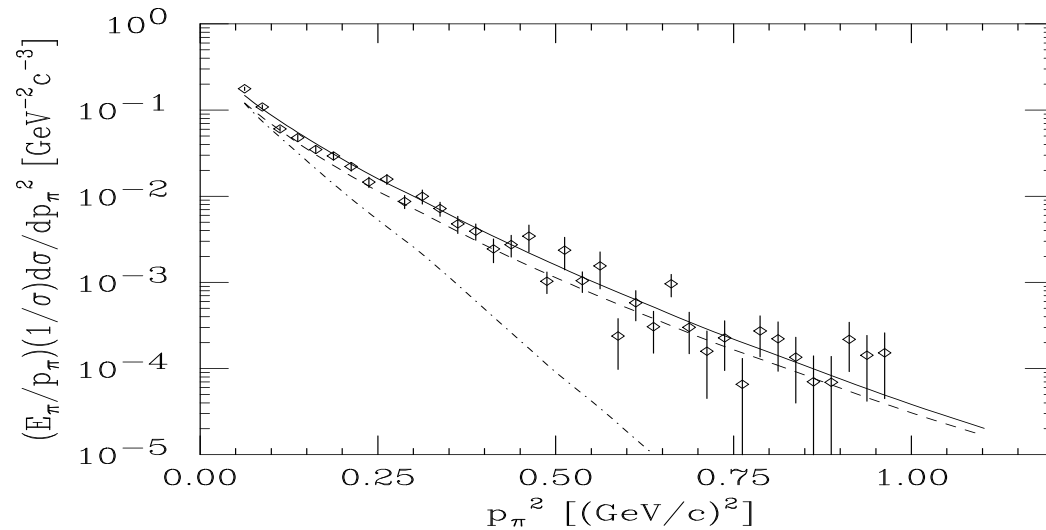


Figure 11: Comparison of our calculations (solid line) with the NOMAD data (full circles) and the MC calculation (open circles)

MC calculation and NOMAD data

(*P.Astier, et al., (NOMAD Coll.) Nucl.Phys..B609,255 (2001)*)

Our calculations

(*O.Benhar,S.Fantoni, G.L., Eur.Phys.J.A7,415 (2000);*

O.Benhar,S.Fantoni, G.L., U.Sukhatme, V.V.Uzhinsky,

Eur.Phys.J.A19,147 (2004))

SUMMARY (Low energies)

- I. The FLT can be applied to compute total cross sections for neutrino scattering off heavy nuclei at low neutrino energies.*
- II. The obtained cross sections do not contradict to other calculations within different nuclear models.*
- III. The suggested approach is much simple in comparing to other models.*
- IV. The cross sections obtained within the FLT are different from the results obtained within the Fermi gas approximation in a factor 2.5-3 at $E_\nu \leq 5 - 6\text{MeV}$.*
- V. At higher energies such difference becomes smaller.*
- VI. The suggested approach can be applied to compute the background from solar neutrinos interacting within a detector.*

SUMMARY (High energies)

I. *The standard QCD model analyzing the multiple hadron production in lepton-proton interactions has to be corrected at $Q^2 < 10(\text{GeV}/c)^2$.*

II. *The non perturbative corrections can be included applying the $1/N$ expansion in QCD.*

III. *The inclusion of cylinder graphs or one-Pomeron exchange diagrams leads to satisfactory description of existing experimental data.*

IV. *Application of suggested approach and assuming an existence of non nucleon degrees of freedom in nuclei allows us to describe the NOMAD data on pion production in backward semi-sphere in $\nu - A$ semi-inclusive processes.*

V. *The suggested approach can be applied to analyze experiments like OPERA performed at the LNGS.*